

Simple de Sitter Solutions

Eva Silverstein

Department of Physics and SLAC
Stanford University
Stanford, CA 94305, USA

We present a framework for de Sitter model building in type IIA string theory, illustrated with specific examples. We find metastable dS minima of the potential for moduli obtained from a compactification on a product of two Nil three-manifolds (which have negative scalar curvature) combined with orientifolds, branes, fractional Chern-Simons forms, and fluxes. As a discrete quantum number is taken large, the curvature, field strengths, inverse volume, and four dimensional string coupling become parametrically small, and the de Sitter Hubble scale can be tuned parametrically smaller than the scales of the moduli, KK, and winding mode masses. A subtle point in the construction is that although the curvature remains consistently weak, the circle fibers of the nilmanifolds become very small in this limit (though this is avoided in illustrative solutions at modest values of the parameters). In the simplest version of the construction, the heaviest moduli masses are parametrically of the same order as the lightest KK and winding masses. However, we provide a method for separating these marginally overlapping scales, and more generally the underlying supersymmetry of the model protects against large corrections to the low-energy moduli potential.

December 2007

1. Introduction

Cosmological observations and conceptual questions of quantum gravity motivate string-theoretic models of de Sitter space and inflation (for reviews, see [1,2]). Several general classes of constructions of metastable dS_4 have been outlined in different corners of string theory with various scales of supersymmetry breaking [3][4,5] [6]. These followed earlier work on flux stabilization such as [7,8,9,10,11] and the original realization [12,13] that string theory produces a finely-spaced discretuum of flux contributions to the moduli potential.

Simple and explicit models of compactification down to AdS_4 have been found using this general framework¹ in type IIB string theory [18,19,20] and IIA [14,16,21,22,15,23]. The latter set are particularly appealing, as they make use of power law effects in the string coupling and inverse radii to stabilize all the moduli in some examples [14]. A number of works (e.g. [24]) have been developing a systematic treatment of the consistency conditions for the ingredients involved in this class of compactifications (as well as more general candidate models involving “nongeometric fluxes” such as [25,26,27]). Although some of the previously outlined de Sitter constructions use only such power law forces [3,6], none attain the explicit simplicity of the known AdS_4 models.

In this work, we obtain a reasonably simple and explicit set of metastable dS_4 minima of the moduli potential by combining the most basic features of [6] and [14,16,22,15,23,24], using classic classical results of [28]. From [6] we take the strategy of using negative scalar curvature as a leading positive term in the potential, but in this case we use a simpler compactification (an orientifold of a product of two Nil three-manifolds). The curvature energy competes against – without overwhelming – the contributions from orientifolds, branes, and RR fluxes in the subsequent orders in the expansion in the string coupling g_s .

¹ Some authors, notably T. Banks, have questioned the use of the effective field theories descending from string theory in backgrounds such as de Sitter or anti de Sitter space which are, globally, infinitely far away from the flat spacetime or linear dilaton backgrounds in which the effective theories were originally derived. Moreover, as in [14,15,16], we will consider massive IIA supergravity, which does not have an exactly flat spacetime solution and so has not been derived precisely from a string S-matrix in any background (see [17] for an exploration of duality in this context). However the energy densities in our solution will be small away from defects (whose tensions and charges are well understood in weakly curved spacetime), and we regard it as a conservative working hypothesis that the effective theory applies in this regime.

Nil manifolds are twisted versions of tori, allowing us to fairly straightforwardly generalize the mechanism employed in [14] in the toroidal orientifold case.

In addition to providing positive potential energy, the geometry – and corresponding topology – of our compactification manifold automatically plays two other very useful roles. First, in contrast to the zero curvature case, the curvature yields positive mass squared for some angular metric deformations, an effect which can be understood from the reduced isometry group of the compactification (which corresponds to a reduced number of massless vector bosons) [28]. Secondly, the topology of the Nilmanifolds support fractional Wilson lines and corresponding fractional Chern-Simons invariants, which provide useful small coefficients of the corresponding terms in our moduli potential. This feature of our construction is similar to the strategy applied earlier to heterotic Calabi-Yau compactifications in [29]. In the present case of compactifications on nilmanifolds, there is an infinite sequence of spaces with a finer and finer discretuum of fractional Chern-Simons terms.

The topology also supports new sectors of wrapped branes. In order to introduce enough perturbative competing forces to obtain de Sitter solutions, we introduce KK monopoles (which are five-branes filling space and wrapping a two-cycle in the compactification). These branes break the supersymmetry at a high but controllable scale: the supersymmetry breaking scale is at the lowest of the KK mass scales in the geometry. For this reason – and also to exhibit the basic physical forces in the problem – we work directly with the scalar potential in four dimensions. (See [30] for SUSY-breaking orbifolds of the AdS_4 models of [14] which also break supersymmetry at a high scale in a controlled way.)²

It would be interesting to investigate the possibility of a similar mechanism preserving lower-energy SUSY, and we will mention some ideas in this direction. In any case, it is perhaps worth emphasizing that after supersymmetry breaking, the methods for gaining control of solutions in the effective theory are essentially the same at different scales of supersymmetry breaking: one requires control over perturbative quantum and α' corrections via a well-defined approximation scheme in which the forces used to stabilize the moduli are the dominant ones. The non-renormalization theorems of supersymmetry, while helpful in restricting the set of corrections to compute, can at the same time complicate moduli stabilization by preventing useful contributions to the moduli potential in the first place. For this reason we obtain our explicit solutions most easily without imposing low-energy SUSY, although the simplifications of ten-dimensional SUSY will play a useful role.

² For recent discussions of the SUSY packaging of the effective action from compactifications on nilmanifolds with various fluxes, see e.g. [24,31].

We will exhibit sequences of solutions which have parametrically small curvature, flux densities, and string coupling as we take an integer quantum number M to be large. Moreover, in the solutions there is a tuneably small ratio between the de Sitter Hubble scale H and the masses of the scalar fields.

However, there is a subtlety in our construction. In the large M limit, although the curvature becomes parametrically weaker we will have a small radius $L_x \sqrt{\alpha'} \ll \sqrt{\alpha'}$ along two directions of the solution (the circle fibers of the nilmanifolds). Relatedly, in the simplest version of the construction, the heaviest moduli masses end up of the same order as the lightest KK and winding masses. Nonetheless, we will exhibit a representative numerical solution at M of order 10, for which L_x is not substringy and for which the corrections are expected to be small since the couplings and curvatures are weak. Finally, in the parametric $M \gg 1$ limit we will suggest a more elaborate method to push the (otherwise marginally overlapping) moduli, winding, and KK scales apart from each other (by introducing extra NS5-branes which locally reduce the inverse string coupling and hence the KK fivebrane tensions). In any case, this small- L_x limit remains a regime of low curvature and approximate supersymmetry as we will explain at the relevant points.

A nilmanifold by itself could simply be T-dualized along the circle fiber directions to a torus with Neveu-Schwarz H flux [32], but our construction involves other ingredients such as nontrivial Neveu-Schwarz B fields and H flux and we will stay in our original T-duality frame for convenience. It is interesting to note that in the models of [14], the moduli masses were of order the curvature scale of the AdS_4 . Here, this problem is avoided, with the moduli masses ending up well above the Hubble scale of our de Sitter – but in the simplest parametric limit they bump up against the next higher scale in the problem, the mass scale of the KK modes.

In model building in general and moduli stabilization in particular, it is important to separate the “forest” (the general mechanisms) from the “trees” (the idiosyncrasies of a given construction). One of the general lessons of the present work – obtained via a simple way of organizing the analysis – is that the AdS_4 models [14,16,22,15,23] and the like admit “uplifting” terms in their potential from a combination of negative scalar curvature and branes. At the level of the overall volume and string coupling, the first point was also made recently in the interesting work [33] which we received as this paper was in preparation (and see [34] for an investigation of using quantum effects to obtain dS solutions). A second general lesson is that the topology of spaces of negative scalar curvature naturally

supports fractional Wilson lines and fractional Chern-Simons invariants, which yield useful small coefficients in the potential.

Although a generic cosmological solution in the landscape is quite complicated – a fact that may be crucial for modeling the observed scale of the dark energy as proposed in [12] – specific models are useful. If completely explicit, such constructions remove the possibility of a conspiracy working against the genericity arguments employed in the general proposals. Conversely, their details expose limitations to tuneability of parameters in specific contexts. In this spirit, recently a clean no-go result for inflation in IIA Calabi-Yau compactifications (with a subset of the possible orientifolds, fluxes, and branes) was given in [33]. In constructing de Sitter in IIA, we were naturally led to ingredients going beyond the assumptions of [33], and together these results make it possible to focus on an appropriate set of degrees of freedom to obtain accelerated expansion in IIA.

In particular – and this is one of the main motivations of the present work – explicit constructions facilitate concrete study of the question of what microphysical degrees of freedom are required to formulate cosmological spacetimes, perhaps in the same way that concrete black brane solutions facilitated the development of black hole state counting and the AdS/CFT correspondence. Nilmanifolds, like hyperbolic spaces, play a central role in geometric group theory [35], and hence compactifications on them may be of further conceptual interest (either at the perturbative level [36,37] or holographically [38]).

The organization of the paper is as follows. In §2, after recording our ten dimensional action and conventions we provide a convenient way of organizing the problem of checking for de Sitter minima of the moduli potential. In §3, we describe a particular class of models on nilmanifolds in detail. We first describe each ingredient and its contribution to the four dimensional potential, noting subtle features as they arise. Next, we demonstrate the stabilization of the coupling and volumes analytically and numerically, noting the behavior of the relevant scales in the parametric limit of interest and suggesting a more involved setup which separates the scales further. We separately analyze the angular moduli, showing how curvature and the other ingredients source them and can be arranged to lift them; we also note various orbifold variants of the model which could be used to project out many of the angular moduli. In §4, we outline a generalization to simple de Sitter solutions on Sol manifolds. Finally, we conclude in §5. Illustrated step-by-step instructions are included at the end.

2. Preliminaries

2.1. Action and conventions

We will follow some of the conventions of [14] (which itself followed [15]); for example our RR fields satisfy $C_{RR} = C_{RR}^{Polch}/\sqrt{2}$ relative to the conventions in [39]. We will start from the ten-dimensional limit of type IIA string theory, for which the action contains kinetic terms

$$\mathcal{S}_{kin} = \frac{1}{2\kappa^2} \int d^{10}x \sqrt{-G} \left(e^{-2\phi} \left(\mathcal{R} + 4(\partial_\mu \phi)^2 - \frac{1}{2}|\tilde{H}|^2 \right) - \left(|\tilde{F}_2|^2 + |\tilde{F}_4|^2 + m_0^2 \right) \right) \quad (2.1)$$

(with $|F_p|^2 = F_{\mu_1 \dots \mu_p} F^{\mu_1 \dots \mu_p} / p!$ and $2\kappa^2 = (2\pi)^7 (\alpha')^4$). Here the full field strengths include Chern-Simons terms

$$\begin{aligned} \tilde{H}_3 &= dB + H_3^{bg} \\ \tilde{F}_2 &= dC_1 + m_0 B \\ \tilde{F}_4 &= dC_3 - C_1 \wedge H_3 - \frac{m_0}{2} B \wedge B \end{aligned} \quad (2.2)$$

where the RR zero-form field strength F_0 is denoted m_0 as in [14]. We will need zero-form, 3-form and 6-form fluxes. As discussed in [14], the flux quantization conditions can be written as

$$m_0 = \frac{f_0}{2\sqrt{2\pi\sqrt{\alpha'}}} \quad p = (2\pi)^2 \alpha' h_3 \quad K = \frac{1}{\sqrt{2}} (2\pi)^5 (\alpha')^{5/2} f_6 \quad (2.3)$$

where f_0, f_6 and h_3 are integers. Here $H = p\omega_3$ is the NS flux through a 3-cycle $\hat{\Sigma}_3$, with ω_3 normalized such that $\int_{\hat{\Sigma}_3} \omega_3 = 1$, and $F_6 = K\omega_6$ where the integral of ω_6 over the compactification manifold is equal to one. We will also make use of fractional Wilson lines descending from B and the corresponding fractional Chern-Simons invariants following from (2.2) in the presence of nontrivial m_0 flux. The B field is normalized in the conventional way so that it appears in the worldsheet path integral via the factor $Exp[\frac{i}{2\pi\alpha'} \int B]$. Its periodicity is

$$\int_{\Sigma_2} B = \int_{\Sigma_2} B + (2\pi)^2 n \quad (2.4)$$

for integer n .

In the next section, the curvature and flux terms in (2.1) as well as orientifold planes and KK5-branes will yield contributions to the potential energy \mathcal{U} in four dimensions upon compactification on a \mathbb{Z}_2 orientifold \mathcal{N}/\mathbb{Z}_2 of volume $L_0^6 (\alpha')^3 / 2$. (The covering space \mathcal{N} of the orientifold has volume L_0^6 , hence our notation.) We will also mention the possibility

of further orbifolding prescriptions, which would modify the volume, flux quantization, and tadpole cancellation conditions in a way which can be obtained via a straightforward generalization of the unorbifolded case.

As reviewed explicitly in [1], it is most convenient to work in four-dimensional Einstein frame obtained by scaling out of the kinetic terms the dependence on the dynamical scalars. Denoting $e^\phi \equiv g_s e^{\tilde{\phi}}$ and $L^6/2 \equiv (L_0^6/2)e^{6\sigma}$, with $\tilde{\phi}$ and σ fluctuating scalar fields, we change variables to

$$G_{\mu\nu,E}^{(4)} = e^{6\sigma-2\tilde{\phi}} G_{\mu\nu,S}^{(4)} \quad (2.5)$$

where $G_{\mu\nu,S}^{(4)}$ denotes the four-dimensional components of the string-frame metric G appearing in (2.1). The four-dimensional potential energy density in Einstein frame is then given by

$$\mathcal{U} = M_4^4 \frac{e^{4\phi}}{(L^6/2)^2} \mathcal{U}_s. \quad (2.6)$$

where $M_4 \sim \frac{L_0^3}{g_s \sqrt{2\alpha'}}$ is the four dimensional Planck mass scale and $\mathcal{U}_s \equiv -\frac{1}{2(\alpha')^2} \int_{\mathcal{N}/\mathbb{Z}_2} e^{-2\phi} \mathcal{R} + \dots$ is the potential energy in string frame.

2.2. Structure of the Potential

Starting from a type II perturbative string limit and defining the 4d coupling

$$g = \frac{e^\phi}{(L^3/\sqrt{2})} \quad (2.7)$$

the moduli potential in four-dimensional Einstein frame has the form

$$\mathcal{U} = M_4^4 g^2 (a - bg + cg^2) + \dots \quad (2.8)$$

where a, b, c depend on other moduli σ_I . Taking the case with $a, b, c > 0$ at the minimum in the σ_I directions, and solving the quadratic equation obtained from imposing $g \partial_g \mathcal{U} = 0$ reveals [3] that at fixed a, b, c , a positive energy solution obtains if

$$1 < \frac{4ac}{b^2} < \frac{9}{8}. \quad (2.9)$$

Violating the lower constraint yields AdS rather than dS, while violating the upper constraint removes the local minimum in the potential in the g direction; i.e. for $4ac/b^2 = 9/8$ there is an inflection point in the potential. Parameterizing the third coefficient in (2.8)

by $c = \frac{b^2}{4a}(1 + \delta)$, the range (2.9) corresponds to $0 < \delta < 1/8$ and the potential takes the form

$$\frac{\mathcal{U}}{M_4^4} = g^2 a \left(1 - \frac{b}{2a}g\right)^2 + \delta \frac{b^2}{4a}g^4 \quad (2.10)$$

Because the algebra involved in minimizing the potential can get somewhat complicated in practice, it proves useful to organize the problem by first minimizing δ at a value slightly above zero and then showing that there is a nearby minimum of \mathcal{U} itself.

In particular, we will exhibit a compactification with the following property. Minimizing the quantity $4ac/b^2$ as a function of the other moduli σ_I yields a value in the range (2.9); we will explicitly use discrete quantum numbers to tune the minimal value of $4ac/b^2$ to be close to but slightly greater than 1. That is, we start in a configuration $\sigma_I = \sigma_{I,0}$ minimizing δ , with $\delta_0 \approx 0$.

If we had $\delta_0 = 0$, then \mathcal{U} would be minimized in the g direction at $g_0 = 2a_0/b_0$, and it is immediately clear that the potential (2.10) would rise quadratically in each direction in field space away from the configuration $\sigma_{I0}, g_0 = 2a_0/b_0$. That is, minimizing δ would also minimize \mathcal{U} (at fixed $g = g_0$). For δ_0 tuned to be small but nonzero, there is still a local de Sitter minimum of the potential which is close to σ_{I0}, g_0 in field space, as can be seen as follows. For $g = g_0 = 2a_0/b_0$ and $\sigma_I = \sigma_{I0}$ (the values minimizing δ), with small positive δ_0 , the potential is of order

$$\mathcal{U} \sim \delta_0 \bar{\mathcal{U}} \quad (2.11)$$

and there is a small tadpole

$$\frac{\partial \mathcal{U}}{\partial \sigma_I} \sim \delta_0 \bar{\mathcal{U}} \quad (2.12)$$

where $\bar{\mathcal{U}}$ is of the same order as the individual terms in the potential $a_0 g_0^2 \sim b_0 g_0^3 \sim g_0^4 \frac{b_0^2}{4a_0}$. The distance the fields are pushed by this tadpole is small, however, because in this configuration σ_{I0}, g_0 there is also a positive quadratic term which is not suppressed as $\delta_0 \rightarrow 0$:

$$\frac{\partial^2 \mathcal{U}}{\partial \sigma_I^2} \sim \frac{\partial^2 \delta}{\partial \sigma_I^2} \bar{\mathcal{U}} \sim \bar{\mathcal{U}} \quad (2.13)$$

Similar scalings to (2.12)(2.13) apply to the derivatives with respect to the dilaton. The result is that the small tadpoles shift the fields a distance of order δ_0 in field space to a local minimum. At this local minimum, the potential is still of order (2.11) (plus subleading terms of order δ_0^2).

After specifying our model, we will show analytically that there is a minimum δ_0 of δ , which can be tuned close to zero by appropriate choices of discrete quantum numbers,

thus providing a de Sitter minimum of the potential. We will then check our results by numerically exhibiting the corresponding minimum of the potential for the coupling and volumes, for specific values of the discrete quantum numbers.

3. Models on Nilmanifolds

Let us start from type IIA string theory in ten dimensions and consider a compactification on an orientifold of a product $\mathcal{N} \equiv \mathcal{N}_3 \times \tilde{\mathcal{N}}_3$ of two Nil three-manifolds (a.k.a. twisted tori, a.k.a. spatial sections of Bianchi II cosmologies, a.k.a. three-tori with “metric flux”). These manifolds are obtained starting from the noncompact geometry

$$\begin{aligned} \frac{ds^2}{\alpha'} &= L_{u_1}^2 du_1^2 + L_{u_2}^2 du_2^2 + L_x^2 \left(dx + \frac{M}{2} [u_1 du_2 - u_2 du_1] \right)^2 + \\ &+ L_{\tilde{u}_1}^2 d\tilde{u}_1^2 + L_{\tilde{u}_2}^2 d\tilde{u}_2^2 + L_{\tilde{x}}^2 \left(d\tilde{x} + \frac{M}{2} [\tilde{u}_1 d\tilde{u}_2 - \tilde{u}_2 d\tilde{u}_1] \right)^2 \\ &= L_{u_1}^2 \eta_1^2 + L_{u_2}^2 \eta_2^2 + L_x^2 \eta_3^2 + L_{\tilde{u}_1}^2 \tilde{\eta}_1^2 + L_{\tilde{u}_2}^2 \tilde{\eta}_2^2 + L_{\tilde{x}}^2 \tilde{\eta}_3^2 \end{aligned} \quad (3.1)$$

where M is an integer and $\eta_1 = du_1$, $\eta_2 = du_2$, $\eta_3 = dx + \frac{M}{2} [u_1 du_2 - u_2 du_1]$ are one-forms invariant under the Heisenberg group of isometries of the nilgeometry (and similarly for the tilded coordinates). We compactify this space by making identifications on the coordinates by a discrete subgroup of the isometry group generated by elements:

$$\begin{aligned} t_x : (x, u_1, u_2, \tilde{x}, \tilde{u}_1, \tilde{u}_2) &\rightarrow (x + 1, u_1, u_2, \tilde{x}, \tilde{u}_1, \tilde{u}_2) \\ t_1 : (x, u_1, u_2, \tilde{x}, \tilde{u}_1, \tilde{u}_2) &\rightarrow (x - \frac{M}{2} u_2, u_1 + 1, u_2, \tilde{x}, \tilde{u}_1, \tilde{u}_2) \\ t_2 : (x, u_1, u_2, \tilde{x}, \tilde{u}_1, \tilde{u}_2) &\rightarrow (x + \frac{M}{2} u_1, u_1, u_2 + 1, \tilde{x}, \tilde{u}_1, \tilde{u}_2) \end{aligned} \quad (3.2)$$

and similarly for the tilded coordinates. The Nil 3-manifold can be described as follows. For each u_1 , there is a torus in the u_2 and $x' \equiv x - \frac{M}{2} u_1 u_2$ directions (under this change of coordinates we have $\eta^3 = dx' + M u_1 du_2$). Moving along the u_1 direction, the complex structure τ of this torus goes from $\tau \rightarrow \tau + M$ as $u_1 \rightarrow u_1 + 1$. The projection by t_{u_1} identifies these equivalent tori. The directions u_1 and u_2 are on the same footing; similar statements apply with the two interchanged and with x' replaced by $x'' \equiv x + \frac{M}{2} u_1 u_2$.

We will orientifold the space by an exchange of the tilded and untilded coordinates combined with an exchange of left and right movers; hence the volume of our compactification will be $L^6/2$ where

$$L^6 = L_x^2 L_{u_1}^2 L_{u_2}^2 \quad (3.3)$$

is the volume of the full compact space \mathcal{N} in string units.

The projections (3.2) generate the fundamental group of \mathcal{N} . They satisfy the relation

$$t_2 t_1 t_2^{-1} t_1^{-1} = t_x^M \quad (3.4)$$

The first homology is given by the abelianization of the fundamental group, obtained by setting all commutators to the identity. For each nil three-manifold this is $\mathbb{Z}^2 \times \mathbb{Z}_M$. The last factor comes from cycles introduced by the projections t_x^m with $m < M$ (since for $m < M$ these elements are not commutators in the fundamental group, and hence are not set to the identity by the abelianization). The nilmanifold with $M \neq 1$ is a freely acting \mathbb{Z}_M orbifold of the a nilmanifold with $M = 1$, obtained by projecting by translations along the x direction. The smaller cohomology group than that of the 3-torus arises because of the relation

$$d\eta_3 = M\eta_1 \wedge \eta_2, \quad (3.5)$$

which means η_3 is not closed and $\eta_1 \wedge \eta_2$ is exact, reducing by one the dimension of H^1 and H^2 . As a result, there are fewer continuous moduli from NS and RR gauge potentials on nilmanifold compactifications as compared to tori, and there are additional vacua corresponding to discrete Wilson lines which we will employ.

The compact nilmanifold also has a reduced isometry group: upon compactification (3.2), the nilmanifold retains only the $U(1)$ isometry corresponding to continuous shifts of x , in contrast to the $U(1)^3$ isometry group of T^3 . As mentioned above, this will help lift some of the scalar degrees of freedom which are eaten in the generalized Higgs mechanism explained in [28].³

We have chosen a symmetric configuration (3.1) to expand around. This renders the analysis simpler since enhanced symmetry points are automatically extrema of the full effective potential in symmetry-breaking directions. Of course we must lift all the light scalar fields including the symmetry-breaking approximate moduli of the metric as well as L_{u_1}, L_{u_2}, L_x and the string coupling g_s . Throughout the construction, for simplicity we will maintain a symmetry between (x, u_1, u_2) and $(\tilde{x}, \tilde{u}_1, \tilde{u}_2)$, a symmetry which will be enforced by an orientifold projection.

³ This fact that a non-Ricci-flat compactification introduces fewer moduli than its Ricci-flat counterpart is an example of a more general phenomenon; hyperbolic spaces of dimension greater than two are famously rigid, there being no continuous deformations of the isometry groups used to compactify the space by projection from the hyperboloid.

The scalar curvature of $\mathcal{N} = \mathcal{N}_3 \times \tilde{\mathcal{N}}_3$ is

$$\mathcal{R} = -\frac{L_x^2 M^2}{\alpha' L_u^4} \quad (3.6)$$

where $L_u^4 \equiv L_{u_1}^2 L_{u_2}^2$. This contributes a positive term in the four-dimensional Einstein frame potential energy (2.6)

$$\mathcal{U}_{\mathcal{R}} = \frac{M_4^4}{2} \frac{L_x^2 M^2}{L_u^4} \frac{e^{2\phi}}{(L^6/2)} = \frac{M_4^4}{2} \frac{e^{2\phi} M^2}{(L^6/2)} \frac{L_x^4}{L^6} = M_4^4 \frac{g^2 L_x^4 M^2}{2L^6} \quad (3.7)$$

descending from the $10d$ Einstein-Hilbert action, where we used (3.3) and the above definition of g (2.7). An important feature of (3.6)(3.7) is the fact that for $ML_x/L_{u_1} \ll 1$ and $ML_x/L_{u_2} \ll 1$, the inverse curvature radius is smaller than the KK scales $1/L_x, 1/L_{u_1}$, and $1/L_{u_2}$. This is related to the fact that the nilmanifold is T-dual, along the x direction, to a T^3 with NS three-form flux – a system for which moduli masses are below the KK scale of the T^3 [32]. As the x circle *shrinks* ($L_x \rightarrow 0$) the curvature becomes *weaker*.

In our final solution, the curvature and flux densities will be small⁴ but M will be large enough that $L_x^2 M \sim 1$ parametrically as $M \rightarrow \infty$. At modest values of M (e.g. $M \sim 10$), we will find numerically that L_x can be slightly greater than string scale. In the parametric limit at small L_x , one can consider T-dualizing to obtain a large circle, but we will continue to describe the system in the original T-duality frame. One reason for this is that our solution will involve discrete Wilson lines from the NS B field as well as NS flux, which complicate the T-duality transformation. Although L_x gets small in this limit, the winding modes will remain parametrically at least as heavy as the moduli and the lightest KK modes in all versions of the constructions.

In the most symmetric case where the lightest KK modes are not be parametrically heavier than the heaviest moduli, but we will still analyze and stabilize the moduli fields separately from the KK modes, for two reasons. First, the KK (and winding) modes – treated separately themselves – are massive (those with winding or momentum on the $T_{x,\tilde{x}}^2$ exhibiting interesting Landau level degeneracies on the nilmanifold with H flux [37]). This together with our analysis of the moduli masses will establish that the diagonal blocks in the moduli and KK mass matrices are positive. The main remaining question in this version of the construction is then whether large off-diagonal terms in the mass matrix

⁴ For substantial recent progress controlling worldsheet theories with substantial curvature and H flux, see [40].

could arise. As we will see after assembling our ingredients, the symmetries of the problem help suppress mixing between the lightest KK modes and the heaviest moduli, suggesting that this need not happen (though we have not analyzed this combined problem in nearly as much detail as the moduli themselves). Because of this uncertainty, we will also suggest a generalization of the model with an extra ingredient which allows us to push the marginally overlapping scales apart.

This contribution (3.6)(3.7) pertains to the diagonal metric (3.1); the curvature will also depend on a subset of the off-diagonal deformations, lifting them in a way originally computed in [28]. The approximate moduli of nilmanifolds were laid out in a form respecting the symmetry structure of the theory in [28]. They consist of metric deformations, deformations of the NS NS two-form potential, and RR axions.

Let us start with the metric moduli. The metric modes are

$$\Delta ds^2 = \mathcal{G}_{IJ}\eta^I\eta^J + \mathcal{G}_{\tilde{I}\tilde{J}}\tilde{\eta}^{\tilde{I}}\tilde{\eta}^{\tilde{J}} + \mathcal{G}_{I\tilde{J}}\eta^I\tilde{\eta}^{\tilde{J}} \quad (3.8)$$

with $G_{IJ} \equiv G_{\tilde{I}\tilde{J}}$ enforced by an orientifold action we will introduce below. Of these, \mathcal{G}_{xu_i} , $\mathcal{G}_{\tilde{x}\tilde{u}_i}$, and $\mathcal{G}_{u_i\tilde{u}_j}$ are lifted in the Higgs mechanism explained in [28]. As discussed above, in contrast to the $U(1)^3$ isometry group of a T^3 , only one $U(1)$ isometry $x \rightarrow x + \lambda$ survives from each Nil three-manifold. In compactification, isometries yields lower-dimensional gauge bosons from off-diagonal metric modes. The would-be gauge bosons corresponding to the broken $U(1)^2$ still exist in the present case of a twisted three-torus, but in a Higgsed phase.

Consider now the $\mathcal{G}_{x\tilde{x}}$ mode. Dimensionally reducing first on one of the nilmanifold factors, say \mathcal{N}_3 , this is a component A_x^{met} of a $U(1)$ gauge boson A_μ^{met} arising from the continuous isometry of the metric in the x direction. The Wilson line of this gauge boson around the \tilde{x} direction is constrained by the relation (3.4):

$$(e^{i \oint_{\gamma_{\tilde{x}}} A^{met}})^M = 1 \Rightarrow \oint_{\gamma_{\tilde{x}}} A^{met} = 2\pi \frac{q'}{M} \quad q' \in \mathbb{Z}. \quad (3.9)$$

In particular, there is not a continuous Wilson line degree of freedom associated with this mode: the mode $A_{\tilde{3}}\tilde{\eta}^3$ is massive, as can be seen from the formulae in [28] (where the discrete Wilson line degree of freedom is not described directly). A similar statement holds also for the NS B field. We will find that discrete Wilson line degrees of freedom are very useful in our setup, and will describe them and their effects in more detail below.

Among the angular moduli, this curvature potential leaves unfixed the modes $\mathcal{G}_{u_1 u_2}, \mathcal{G}_{\tilde{u}_1 \tilde{u}_2}, \mathcal{G}_{u^i \tilde{x}} = \mathcal{G}_{x \tilde{u}^i}$ ($i = 1, 2$). The modes which are lifted by the curvature can also get contributions from other terms in the moduli potential, as we will discuss in analyzing the angular moduli below in §3.7. In addition, we must stabilize the diagonal moduli $\mathcal{G}_{II} = \mathcal{G}_{\tilde{I}\tilde{I}}$, equivalently $L_x^2, L_{u_1}^2, L_{u_2}^2$ (one of which can be traded for the overall volume mode (3.3), which is a runaway direction in field space).

Some of the B, C_1, C_3 , and C_5 fields will be lifted by a combination of the \tilde{F}^2 terms from (2.1)(2.2) and by the orientifold projection (and in more general examples, orbifold projections). Others will be unsourced by the leading terms in the potential, and be fixed by higher order, lower scale effects.

3.1. Orientifold and fluxes

In order to obtain a metastable solution, we will require a negative term in the potential, at an intermediate order in the expansion about weak coupling and large volume, since the potential energy in four dimensional Einstein frame decays to zero at weak coupling and low curvature [1]. To this end, introduce an O6-plane as in [14], as follows.⁵ Mod out the worldsheet sigma model by an exchange of tilded and untilded embedding coordinates (and Fermi partners) combined with an exchange of left and right movers:

$$\Omega : (x, u_j, \tilde{x}, \tilde{u}_k) \leftrightarrow (\tilde{x}, \tilde{u}_j, x, u_k) \quad L \leftrightarrow R \quad (-1)^{F_L} \quad (3.10)$$

As reviewed in [14], under the orientifold transformation, B, C_1 , and C_5 are odd. Geometrically, (3.10) introduces on O6 plane wrapped on the 3-cycle traced out by the fixed point locus $(x, u_j) = (\tilde{x}, \tilde{u}_j)$. The negative tension of the O6 plane leads to the potential energy contribution

$$\mathcal{U}_O = -\kappa^2 (2\mu_6) \frac{g_s^3}{(L^6/2)^2} (Vol_{O6}) = -2^3 \pi g^3 \quad (3.11)$$

⁵ There are other, discretely distinct, options for defining the space group of the orientifold. Another example would be to mod by the same orientifold action, but – in compactifying the original space via the discrete isometries (3.2) – to project only by the group generated by elements of the form $t_y t_{\tilde{y}}$ and $t_y t_{\tilde{y}}^{-1}$. This would yield 2^3 different O6-planes from the fixed points of the group action, on a space of volume $4L^6$. We expect similar results for all these cases, but the detailed factors in the potential would differ in different examples.

where μ_6 is the D6-brane tension (equal to π/κ^2) and Vol_{O6} is the volume of the cycle wrapped by the O6-plane (which in our case is $(\sqrt{2})^3 L^3$).

We must cancel the O6-plane's charge. Following [14], we can use fluxes to cancel the tadpole for C_7 . The O6-plane constitutes a localized source of F_2 within the 3-cycles dual to the cohomology classes $\eta^1 \wedge \eta^2 \wedge \eta^3 - \tilde{\eta}^1 \wedge \tilde{\eta}^2 \wedge \tilde{\eta}^3$, $\eta^1 \wedge \tilde{\eta}^2 \wedge \eta^3 - \tilde{\eta}^1 \wedge \eta^2 \wedge \tilde{\eta}^3$, and $\tilde{\eta}^1 \wedge \eta^2 \wedge \eta^3 - \eta^1 \wedge \tilde{\eta}^2 \wedge \tilde{\eta}^3$. The tadpole cancellation condition is that for each 3-cycle Σ_3 ,

$$m_0 \int_{\Sigma_3} H = -2\sqrt{2}\mu_6\kappa^2 n_{O6} \quad (3.12)$$

where n_{O6} is the net number of O6-planes sitting at points in Σ_3 . Writing

$$\begin{aligned} H \equiv & p_1 \left(\eta^1 \wedge \eta^2 \wedge \eta^3 - \tilde{\eta}^1 \wedge \tilde{\eta}^2 \wedge \tilde{\eta}^3 \right) + p_2 \left(\tilde{\eta}^1 \wedge \eta^2 \wedge \tilde{\eta}^3 - \eta^1 \wedge \tilde{\eta}^2 \wedge \eta^3 \right) \\ & + p_3 \left(\eta^1 \wedge \tilde{\eta}^2 \wedge \tilde{\eta}^3 - \tilde{\eta}^1 \wedge \eta^2 \wedge \eta^3 \right) \end{aligned} \quad (3.13)$$

where $p_i \equiv -h_{3,i}(2\pi)^2\alpha'$ (c.f. (2.3)), and imposing (3.12) (with $n_{O6} = 1$ O6-planes passing through each cycle) sets $f_0 h_{3,i} = 2$. We therefore take $h_{3i} = h_3$ for $i = 1, \dots, 3$, with

$$f_0 = 1 \quad h_3 = 2 \quad (3.14)$$

Note that this satisfies the flux quantization condition on both the covering space and the orientifold itself. In evaluating (3.12), we took into account the fact that each of the 3-cycles is halved in volume by the action of the orientifold (c.f. [41]).

The O6-plane, H_3 flux, and F_0 flux together contribute the following terms to the four-dimensional effective potential in 4d Einstein frame (2.6):

$$\mathcal{U}_{OHm_0} = M_4^4 \left(\frac{3p^2 g^2}{2(\alpha')^2 L^6} - 2\sqrt{\frac{2}{\alpha'}} |m_0 p| g^3 + \frac{\alpha' m_0^2 g^4 L^6}{4} \right) \quad (3.15)$$

We will also include six-form flux

$$F_6 = K \eta^1 \wedge \eta^2 \wedge \eta^3 \wedge \tilde{\eta}^1 \wedge \tilde{\eta}^2 \wedge \tilde{\eta}^3 \quad (3.16)$$

where $K = f_6(2\pi)^5(\alpha')^{5/2}/\sqrt{2}$ in terms of the integer flux quantum number f_6 . This leads to the following term in the moduli potential:

$$\mathcal{U}_{F_6} = M_4^4 g^4 \frac{K^2}{4L^6(\alpha')^5} \quad (3.17)$$

3.2. Fractional Chern-Simons invariants

The $\mathbb{Z}_M \times \mathbb{Z}_M$ homology cycles described above yield a set of discrete Wilson line vacua, and corresponding fractional Chern-Simons invariants which we will use to obtain contributions to the effective potential with tuneably small coefficients, as follows.⁶ As discussed in §2.1, the effective action contains a term $-\frac{1}{2\kappa^2} \int d^{10}x |\tilde{F}_2|^2$ where

$$\tilde{F}_2 = dC_{(1)} + m_0 B \quad (3.18)$$

with m_0 the RR 0-form flux (2.3) and B a Neveu-Schwarz two-form potential. In our background solution, B will be flat (note that we separated the H flux from dB in (2.2) as reviewed in [14]). Its fractional Wilson line vacua lead to fractional Chern-Simons forms $m_0 B$, which will provide useful small coefficients in the moduli potential.

In general, a manifold with nontrivial fundamental group π_1 , can support discrete Wilson line vacua of gauge fields of a gauge group G – flat connections with nontrivial holonomy around non-contractible cycles. As reviewed in [42], they correspond to homomorphisms from π_1 into G , since Wilson lines $U_\gamma = P \exp(i \int_\gamma A)$ must satisfy the group multiplication law $U_{\gamma\gamma'} = U_\gamma U_{\gamma'}$. Since G is abelian in our case, the only discrete Wilson lines arise from closed paths which are nontrivial in homology (which is the abelianization of the fundamental group obtained by setting commutators to 1); elements g of π_1 which are commutators (elements of the form $g = g_1 g_2 g_1^{-1} g_2^{-1}$) have trivial holonomy $U_g = 1$. The first homology group of \mathcal{N}_3 includes the \mathbb{Z}_M factor, represented by the closed paths γ_m introduced by the projection t_x^m with $m < M$ (3.2).

Flat connections for gauge fields on nilmanifolds were derived explicitly in [43]. This construction generalizes to the NS 2-form potential. In a local neighborhood of \mathcal{N} , and for $0 \leq x < 1, 0 \leq \tilde{x} < 1$, we can take B to be

$$B = \frac{q}{M} (2\pi)^2 \alpha' dx \wedge d\tilde{x} + \frac{r}{M} (2\pi)^2 \alpha' (dx \wedge \tilde{\eta}^1 - d\tilde{x} \wedge \eta^1) + (1 \leftrightarrow 2) + \dots \quad (3.19)$$

The first terms of (3.19) contain the B field analogue of discrete Wilson lines. These are projected in by the orientifold action (3.10) (using the fact reviewed in [14] that B has an intrinsic parity under the orientifold, c.f. eqn (2.9) of [14]). The rest indicated by \dots contains the continuous Wilson lines invariant under the orientifold; these are lifted by the

⁶ See [29] for a previous example using fractional CS invariants to help in heterotic moduli stabilization.

$|\tilde{F}_2|^2$ term in the potential. The form for B in different neighborhoods is then derived by using the transition functions between them as in [43].

The discrete Wilson line terms in (3.19) yield

$$e^{\frac{i}{2\pi\alpha'} \int_{\gamma_x \times \gamma_{\tilde{x}}} B} = e^{2\pi i \frac{q}{M}} \quad e^{\frac{i}{2\pi\alpha'} (\int_{\gamma_x \times \gamma_{\tilde{u}_1}} B - \int_{\gamma_{\tilde{x}} \times \gamma_{u_1}} B)} = e^{2\pi i \frac{2r}{M}} \quad (3.20)$$

and introduces the potential term (from the $|\tilde{F}_2|^2$ and $|\tilde{F}_4|^2$ terms in (2.1))

$$\mathcal{U}_{BWL} = 4\pi^4 M_4^4 m_0^2 \alpha' \left(\frac{q}{M}\right)^2 g^4 \frac{L^6}{L_x^4} + 16\pi^4 M_4^4 m_0^2 \left(\frac{r}{M}\right)^2 \frac{g^4 L^3}{L_x} + 2^8 \pi^8 M_4^4 m_0^2 \left(\frac{r}{M}\right)^4 \frac{g^4}{L_x^2} \quad (3.21)$$

We will ultimately choose q to be of order 1, and can use the ratio r/M to help tune the cosmological constant, as well as to help stabilize some angular moduli (though in that regard, orbifold variants of the construction which remove angular moduli could also project out the terms proportional to r/M – this would be a consistent choice, since as we will see these terms are not crucial for stabilizing the coupling and volume moduli). In writing (3.21), we took $L_1 = L_2$, and will consider other ingredients which respect this symmetry and consistently stabilize the system at this point.

Another way to describe the contribution in the first term of (3.19)(3.20)(3.21) is as an example of discrete torsion [44]. The torsion cycles γ_1 and $\gamma_{\bar{1}}$ in our compactification manifold are obtained by the projection $\mathbb{Z}_M \times \mathbb{Z}_M$ starting from a finite cover (the same space with $M = 1$). With the B field (3.19) turned on, the projection in the \mathbb{Z}_M winding string sectors are modified by the factor $Exp[\frac{i}{2\pi\alpha'} \int B]$ in the worldsheet path integral. This example has the interesting feature that the discrete torsion is not associated with an orbifold singularity, since the projection is freely acting.⁷

3.3. KK5 branes

The nontrivial topology of our compactification manifold can also support wrapped branes. Spacefilling KK monopoles [46] will play a useful role, providing a needed independent “uplifting” term in the potential. These are objects magnetically charged under a linear combination of the $U(1)$ isometries along the x, \tilde{x} directions, and are extended in $4d$ as well as along two internal directions. They are T-dual to $NS5$ -branes [47], and we will refer to them as KK fivebranes.

⁷ For recent studied of discrete torsion, see e.g. [45].

As a specific example, we introduce the following set of KK5-branes. Start with n_K KK fivebranes magnetically charged under $U(1)_x \times U(1)_{\tilde{x}}$ with charges (1,1). (These are T-dual to NS5-branes at points on a circle of radius $\sim 1/L_x$.) Wrap these KK fivebranes along the the transverse direction from $x, \tilde{x} = 0$ to $t_x t_{\tilde{x}}^{-1}$ and from $u_1, \tilde{u}_1 = 0, u_2, \tilde{u}_2 = 0$ to $t_{u_1} t_{\tilde{u}_1} t_{u_2} t_{\tilde{u}_2}$. The latter cycle is subject to ordering ambiguities because of the relation (3.4), but any order will give similar scalings in our moduli potential.⁸

Topological Consistency Conditions and n_K

We must make sure our configuration is consistent with all of the previous ingredients, and we must cancel all the relevant charges within the compactification. The number n_K of KK5-branes may be constrained to be a multiple of M or to be combined with antibranes in order to accomplish this. One reason is charge cancellation. Each of the n_K branes wrap an \mathbb{Z}_M homology cycle. As such, it is not a source for gauge bosons of a continuous gauge symmetry group, so Gauss' law does not directly apply to impose charge cancellation. However, for D-branes in this kind of situation, one does find that K-theory charges must be cancelled in compact manifolds [48], and a similar constraint may arise in the present case. To be safe, we will assume such a condition holds in our present context.

Moreover, in discussing the flat connection for B (3.19)(3.20) above, we used the fact that our compactification manifold is a freely acting $\mathbb{Z}_M \times \mathbb{Z}_M$ orbifold of a nilmanifold, giving torsion 1-cycles $\gamma_1, \gamma_{\tilde{1}}$ in the x, \tilde{x} directions which led to the possibility of discrete torsion (3.20). In the presence of $n_K < M$ KK monopoles, however, strings can only be conserved mod n_K [47], so the \mathbb{Z}_M winding charge for strings wound around these cycles is no longer conserved; the cycles $\gamma_1^{n_K}$ and $\gamma_{\tilde{1}}^{n_K}$ bound 2-cycles. By Stokes' law, in this situation the fractional Wilson line $\frac{1}{2\pi\alpha'} \int_{\gamma_1 \times \gamma_{\tilde{1}}} B$ is quantized in units of $1/n_K$ rather than $1/M$. For n_K a multiple of M , with all the KK5-branes sitting at the same point in their transverse directions, we recover the $\mathbb{Z}_M \times \mathbb{Z}_M$ symmetry and the consequent discrete Wilson line (discrete torsion) taken in (3.19)(3.20).

⁸ As with the other ingredients, there are variants of this configuration which could also be considered, with sets of KK5-brane stretched in various different directions. This is important for example in versions in which one orbifolds the geometry, in which the KK5-brane configuration would need to be invariant under the corresponding symmetry. This can be arranged by using sets of fivebranes respecting the orbifold symmetry, or if necessary adding other sets rotated appropriately relative to the original set. Also, NS fivebranes which play a similar role can be used.

There are in general further topological constraints on combinations of branes and fluxes. A canonical example of this type of consistency condition is that of [49]: Dp-branes with H flux on their worldvolume must have a corresponding number of $D(p-2)$ branes ending on them. This type of condition has been generalized to KK monopoles and NS5-branes in [24]. For KK 5-branes, Villadoro and Zwirner [24] find – using various U-duality arguments – constraints on \tilde{F}_2 flux along a two-cycle consisting of the fiber circle times a one-cycle in the brane worldvolume. In our setup, if $n_K < M$ so that the brane wraps a homologically nontrivial cycle in the x, \tilde{x} directions, then the fractional Chern-Simons invariant $m_0 B$ coming from (3.19) (whose flux quantum number is q/M) is nontrivial. As far as we can tell, more analysis would be required to determine if this leads to an anomaly, and if so whether that anomaly could be cancelled by the addition of other branes.

Because of these subtleties, we will keep track of the n_K dependence but will focus on the cases where n_K is a multiple of M , so that each set of branes is homologous to a single brane wrapping a homologically trivial (but homotopically nontrivial) cycle. This evades both generalized K-theoretic subtleties just listed, and also does not require additional antibranes to cancel the charges (though these could be included).

In this case, the discrete torsion (3.20) is consistent with the KK5-branes, as long as they are placed together. (In the T-dual description, the NS5-branes are arranged symmetrically along the T-dual transverse circle, restoring the \mathbb{Z}_M translation symmetry in that description, but are together in the remaining transverse directions.) This provides another example of discrete torsion helping to stabilize moduli [50], significantly simplifying the problem of stabilizing the 5-brane positions since their relative motion is projected out.

Potential Contribution

For values of $L_u \gg M L_x$ for which the fivebranes are well localized within the transverse u, \tilde{u} and T-dual $x + \tilde{x}$ directions, they are locally supersymmetric. The BPS formulas for the tensions yield the following contribution to the potential \mathcal{U} from n_K such sets of KK 5-branes:

$$\mathcal{U}_{KK5} = M_4^4 2 \sqrt{2 \left(\frac{L_1}{L_2} + \frac{L_2}{L_1} \right)} \frac{n_K}{\sqrt{\eta}} g^2 \frac{L_x^{5/2}}{L^{9/2}} \quad (3.22)$$

where we defined

$$\eta \sim \frac{L_{x-\tilde{x}}}{L_{x+\tilde{x}}} \quad (3.23)$$

This degree of freedom η is related to the angular metric discrete Wilson line degree of freedom describe in (3.9). Starting from a given discrete Wilson line vacuum, varying the continuous modulus $\mathcal{G}_{3\bar{3}}$ by changing the angle γ between the x, \tilde{x} directions (at fixed

volume) changes the lengths of the cycles generated by $t_x t_{\tilde{x}}$ and $t_x t_{\tilde{x}}^{-1}$. More precisely, starting from $q' = 0$, in terms of the angle γ , the ratio η is given by

$$\eta = \sqrt{\frac{1-\gamma}{1+\gamma}} \quad (3.24)$$

In analyzing the angular dependence of our potential terms below in §3.7, we will require their dependence on the (3.23), as well as on other similar angular moduli. In addition to the KK fivebrane contribution, the curvature, O6 and H, F_2, F_4 flux contributions depend on η , reducing to (3.7)(3.11)(3.15) when $\eta \rightarrow 1$. Since the Wilson line is discrete, and the corresponding $\mathcal{G}_{3\bar{3}}$ deformation is massive, the curvature potential acquires a factor of the form $1 + (\eta - \eta_{q'})^2 + \dots$. The negative O6-plane tension acquires a factor of $1/\sqrt{\eta}$, since its length increases when η decreases. The H flux term, which threads the dual cycle to that wrapped by the O6-plane, scales like $1/\eta$, since the flux lives in a larger cycle when η increases. Similarly, the second contribution in (3.21) acquires a factor of $1/\eta$ at small η .

We should make one further comment about the formula (3.22). Since they are all together, our M branes have substantial a throat cross section (the size of each KK5-brane being given by the size L_x of its fiber direction [46]), which with M of them adds up to a size ML_x . We should compare this to the size of the compactification in the directions transverse to them. In our simplest solution, both will be of order $M^{1/2}$, so that the KK fivebrane cores bump up against the size of the compactification. This, along with the ratio of KK to moduli masses, motivates a more elaborate setup separating these scales, and indeed we will ultimately find a method to push the transverse size of the space larger than this. In any case, (3.22) gives a good estimate for the parametric scaling of the KK fivebrane contribution to the effective action. We will return to discuss the angular and fivebrane moduli after addressing the problem of stabilizing the runaway moduli g, L_u, L_x .

3.4. Stabilization of volumes and coupling

Altogether, we have a potential energy for g, L, L_x of the form

$$\begin{aligned} \frac{\mathcal{U}}{M_4^4} &= ag^2 - bg^3 + cg^4 \\ &= g^2 \left(M^2 \frac{L_x^4}{2L^6} + (4n_K) \frac{L_x^{5/2}}{L^{9/2}} + \frac{3p^2}{2(\alpha')^2 L^6} \right) - g^3 \left(2\sqrt{\frac{2}{\alpha'}} |pm_0| \right) \\ &\quad + g^4 \alpha' \left(\frac{m_0^2}{4} L^6 + 4\pi^4 m_0^2 \left(\frac{q}{M} \right)^2 \frac{L^6}{L_x^4} + \left(\frac{r}{M} \right)^2 \frac{16\pi^4 m_0^2 L^3}{L_x} + \left(\frac{r}{M} \right)^4 \frac{2^8 \pi^8 m_0^2}{L_x^2} + \frac{K^2}{4L^6 (\alpha')^6} \right) \end{aligned} \quad (3.25)$$

where we set $L_1/L_2 = 1$, since this is where it is stabilized by (3.22), and where we suppressed dependence on angular moduli to be discussed in the next subsection. Following the discussion in §2, we will proceed to show that for suitable choices of discrete quantum numbers, the quantity

$$\begin{aligned} \frac{4ac}{b^2} &\equiv 1 + \delta(L, L_x) \\ &= \frac{(\alpha')^2}{2p^2m_0^2} \left(M^2 \frac{L_x^4}{2L^6} + (4n_K) \frac{L_x^{5/2}}{L^{9/2}} + \frac{3p^2}{2(\alpha')^2 L^6} \right) \\ &\times \left(\frac{m_0^2}{4} L^6 + 4\pi^4 m_0^2 \left(\frac{q}{M} \right)^2 \frac{L^6}{L_x^4} + \left(\frac{r}{M} \right)^2 \frac{16\pi^4 m_0^2 L^3}{L_x} + \left(\frac{r}{M} \right)^4 \frac{2^8 \pi^8 m_0^2}{L_x^2} + \frac{K^2}{4L^6 (\alpha')^6} \right) \end{aligned} \quad (3.26)$$

has a minimum L_0, L_{x0} in the space of L, L_x which is in the range

$$0 < \delta(L_0, L_{x0}) < \frac{1}{8} \quad (3.27)$$

with $\delta(L_0, L_{x0})$ tuneable to be small.

We need to make sure first of all that the minimum of (3.26) is not above $9/8$, which we can show as follows. First, note that at (3.26) includes a constant term $q^2/(16h_3^2) + 3/16$ (from the $1st \times 2nd$ and $3rd \times 1st$ cross terms, respectively).

Next, consider the two terms (from the $1st \times 1st$ and $3rd \times 2nd$ cross terms, respectively)

$$\frac{\alpha'}{2p^2m_0^2} \left(\frac{M^2 m_0^2 \alpha'}{8} L_x^4 + \frac{6\pi^4 m_0^2 p^2 q^2}{\alpha' L_x^4 M^2} \right) \quad (3.28)$$

The terms (3.28) are minimized at $\frac{q\sqrt{3}}{8h_3}$, with

$$L_x^4 \sim \frac{1}{M^2} \quad (3.29)$$

At fixed L_x , the remaining terms in (3.26) diverge for $L \rightarrow 0$ and for $L \rightarrow \infty$. Thus $4ac/b^2$ has a minimum at finite nonzero values of L, L_x . We will next show that this minimum is tuneable to lie in the range (2.9) as close as desired to the lower limit as $M \rightarrow \infty$, and show that the corrections to our solution are small. We will then verify numerically that a de Sitter minimum of the potential (3.25) exists.

Plugging (3.29) into (3.26), it reduces to $(3/16) + (q^2/(16h_3^2)) + \frac{q\sqrt{3}}{8h_3}$ plus a function of L which diverges as $L \rightarrow 0$ or $L \rightarrow \infty$. The fact that it diverges as $L \rightarrow 0$ for any value of r (including zero) is a consequence of the six-form flux contribution (3.17); in the

absence of this contribution, the minimum of $4ac/b^2$ would be at $L = 0$, with the value $(3/4) + (q^2/(4h_3^2)) + \frac{q\sqrt{3}}{2h_3}$. In the presence of (3.17), the minimum can therefore be tuned to sit within the required range (2.9) by adjusting q , r , M and K .

We will also keep track of factors of n_K/M even though we will consider the case $n_K = M$ for the topological reasons discussed above. The reason for this is that the resulting formulas will make clear that in order to generalize the construction to separate its mass scales further, it would be useful to introduce effects which decrease the contribution of the KK fivebranes. We will ultimately propose to do this (below in §3.6) not by reducing the number n_K of them, but by reducing their tension by placing them in a local region the compactification with an enhanced string coupling.

Now consider the two terms

$$\frac{(4n_K)(\alpha')^2 L_x^{5/2}}{2p^2 m_0^2} \left(\frac{m_0^2}{4} L^{3/2} + \frac{K^2}{4L^{21/2}(\alpha')^6} \right) \quad (3.30)$$

These terms are minimized at a value of order $\frac{n_K}{M}(\frac{K}{M})^{1/4}$ obtained for $L \sim K^{1/6}$. Taking into account that L_x is constrained by (3.28), this term combined with the first term of (3.30) prevents any decay mode with $L \rightarrow \infty$ or $L \rightarrow 0$.

In general all the terms in (3.26) are consistently of the same order at the minimum, with the parametric scalings

$$\left(\frac{n_K}{M} \right) \sim \left(\frac{M}{K} \right)^{1/4} \quad L \sim K^{1/6} \quad L_x \sim \frac{1}{M^{1/2}} \quad (3.31)$$

As emphasized above, L_x ends up small in our parametric limit (3.29). For finite values of M , of course, the results depend on order 1 factors. Below, in §3.8, we will exhibit a numerical local minimum of the potential (3.25) in the g, L, L_x directions for which L_x ends up ~ 2 at $M = 10$.

3.5. Scales and Mixing

Using (3.31) we can now indicate the physical scales of interest in our solution (again assuming that as discussed below in §3.7, the angular moduli stay near the original point (3.1) about which we have expanded). The string coupling $g_s = gL^3$ is

$$g_s \sim \frac{2a}{b} L^3 \sim \frac{1}{L^3} \sim \frac{1}{K^{1/2}} \quad (3.32)$$

Since we have a small cycle L_x , it is interesting to consider the T-dual coupling \hat{g}_s (even though the T-duality affects the ingredients listed above in a somewhat complicated way). This is

$$\hat{g}_s^2 \sim \frac{g_s^2}{L_x^4} \sim \frac{M^2}{K} \sim M \left(\frac{n_K}{M} \right)^4 \quad (3.33)$$

Thus in our simplest setup with $n_K = M$, this is large, but if we can elaborate the model to lower the contribution of the KK fivebrane configuration, this could be small; we will suggest a method for achieving this in the next subsection. Since the theory is approximately supersymmetric, the corrections are parametrically at most of order $\hat{g}_s^2 \mathcal{R} \sim \hat{g}_s^2 / L^6 \sim (M/K)^2 \ll 1$. In the numerical solution of §3.8 at a modest value for M of 10, we will see that both g_s and \hat{g}_s can remain \leq order 1 with small curvatures, leading to suppressed corrections. Even in the parametric $M \rightarrow \infty$ limit of the simplest version of the construction, the estimate just given might be too pessimistic: if in the T-dual model one crossed over into the M theoretic regime with $\hat{g}_s \gg 1$, the corrections to the moduli potential should not grow with increasing size of the eleventh dimension, but should fall off at large radius.

Another scale of interest is the core size of our set of KK 5-branes as compared to the size of the compactification. This is of order $M^{1/2}$, and again will bump up against the size of the compactification $L_u \sim (KM)^{1/4}$ in the simplest version, again motivating increasing the ratio K/M .

Using the fact that the canonically normalized moduli fields are the logs of g_s, L, L_x we have that the scale of moduli masses squared (aside from the residual angular moduli and fivebrane positions to be discussed in the next subsection) is

$$m_{moduli}^2 \sim M_4^2 a g^2 \sim M_4^2 \frac{g^2}{L^6} \sim M_4^2 \frac{1}{K^3} \quad (3.34)$$

The lightest KK modes are those propagating in the u, \tilde{u} directions, which scale like

$$m_{KK}^2 \sim \frac{1}{\alpha' L_u^2} \sim M_4^2 \frac{1}{K^3} \left(\frac{K}{M} \right)^{1/2} \quad (3.35)$$

(where we used the relation $1/\alpha' \sim M_4^2 g_s^2 / L^6 = M_4^2 g^2$ between the string and Planck mass scales).

The masses of the strings wound around the L_x direction are of order

$$m_{winding}^2 \sim \frac{L_x^2}{\alpha'} \sim M_4^2 \frac{1}{MK^2} \quad (3.36)$$

(This includes the effects of the B field (3.19), which adds a term of the same order $-BG^{-1}B$ to the winding mass squared.)

Now let us compare these scales to the Hubble scale H of the de Sitter minimum:

$$\frac{\mathcal{U}_{dS}}{M_4^4} \sim \frac{H^2}{M_4^2} \sim \delta \frac{b^2}{4a} g^4 \sim \delta \frac{1}{K^3} \quad (3.37)$$

Since δ can be tuned to be small by adjusting the precise value of K/M and/or r/M given n_K/M , this gives a hierarchy between the de Sitter curvature scale and the scale of the moduli masses:

$$H^2 \ll m_{moduli}^2 \quad \text{for } \delta \ll 1 \quad (3.38)$$

Of course as one adds small corrections, the precise tune in the discrete quantum numbers which one must do to obtain small δ_0 changes accordingly. As in the realistic context, it would not be possible to explicitly tune arbitrarily finely to cancel all the loop corrections, simply because we do not know the value of these loop corrections.

For $K \sim M$ we do not have a hierarchy between the heaviest moduli and the lightest KK modes, and between the KK modes and the winding modes. Also, the fivebrane cores in this case are of the same order as L_u . In the next subsection, we will elaborate the model to separate these scales.

Before turning to that, in the marginal case let us discuss the question of mixing between the lightest KK modes and the heaviest moduli. First, note that in the nilmanifold by itself, the KK modes do not mix linearly with the moduli fields, to good approximation. This can be seen by constructing the Laplacian on the space, but follows more intuitively from the topology and the physics of the “metric flux”. The lightest KK modes – i.e. those which are not parametrically separated in mass scale from the moduli – have no momentum in the small x, \tilde{x} directions or in the u_1, \tilde{u}_1 directions. Dimensionally reducing on the x and \tilde{x} directions, these KK modes are simply uncharged particles on a torus with Kaluza Klein magnetic flux. Since the particles are uncharged, they have the same spectrum as they would on a two torus, and modes of different momentum do not mix linearly. This property continues to hold classically after the orientifold projection is made, for the standard reason that tree-level amplitudes for untwisted modes are inherited. The KK5-branes do however break the translation invariance in the u_2, \tilde{u}_2 directions, and their effects would need to be included in a full analysis of the moduli+KK dynamics in the $K \sim M$ case.

3.6. Separating the Scales

Rather than including the KK modes in the analysis, it might be simpler to dress up the model so as to push apart these marginally overlapping scales. Of course, additional ingredients used to do this must be introduced in a way which does not destabilize the model. There are several approaches to this; we will suggest one method here.

First, note that the relevant ratio of scales is

$$\frac{m_{KK}}{m_{moduli}} \sim L_x^{1/2} L^{3/2} \sim \frac{K^{1/4}}{M^{1/4}} \quad (3.39)$$

(This quantity also controls the ratio of L_u to the core size of the 5-brane collection, the ratio of lightest winding to lightest momentum masses, and the dual $10d$ string coupling.) In our setup discussed above, the combination of (3.29) and (3.30), combined with the requirement (3.27), bounds the quantity $L_x^{1/2} L^{3/2}$ to be of order M^0 in our parametric limit. However, as discussed above, from (3.31) we see that if the contribution from the KK fivebrane tensions were reduced by some factor $\epsilon < 1$, then the quantity (3.39) would be larger, of order $1/\epsilon$. (To be clear: we will keep $n_K = M$ for the topological reasons discussed above; lowering the tension of the KK fivebranes would feed into the scalings (3.31) as if we had reduced n_K .)

One way to arrange this is to introduce a source of varying string coupling $e^{\phi_{loc}(u,\tilde{u})}$ within the compactification, so that the KK fivebranes (whose tensions scale like $e^{-2\phi_{loc}}$) are reduced when they sit at a position within the compactification with increased string coupling (which minimizes their energy). Our notation ϕ_{loc} here refers to the spatially varying dilaton within the compactification – note that the Einstein frame conversion factor (2.6) involves the ambient $4d$ string coupling e^ϕ , so that the effect of an inhomogeneous dilaton on the potential contribution (3.22) is to rescale it by a factor $e^{-2(\phi_{loc}-\phi)}$.

NS fivebranes provide one source of varying string coupling – $e^{\phi_{loc}}$ increases as one moves toward their cores.⁹ Moreover, in our setup, NS fivebranes which are wrapped on the x, \tilde{x} directions each introduce parametrically less potential energy than the leading terms (3.25):

$$\mathcal{U}_{NS5} \sim g^2 \frac{n_5 L_x^2}{L^6} \quad (3.40)$$

⁹ There are variants of this approach – for D- $(p \geq 4)$ -branes the coupling grows away from the cores of the solutions, suggesting a similar mechanism with the KK fivebranes drawn to positions in between added sets of D-branes.

Specifically, comparing this to the curvature term in (3.25) using (3.28)(3.29) we see that as long as $n_5 \leq M$, adding a set of such NS5-branes provides a term in the potential which is at or below the scale of the above ingredients. At the same time, it provides a varying string coupling within the compactification, which we may be able to use to reduce the KK5-brane contribution to the potential.

Since the KK 5-branes wrap one of the u, \tilde{u} directions, the configuration of interest is one in which they skirt the cores of one or more clouds of NS5-branes as they stretch across them, lowering their energy by passing through regions with lower string coupling. Each NS 5-brane is localized in the u, \tilde{u} directions, but they may be distributed so as to minimize the energy of the whole configuration. We will now estimate whether this effect can be significant in our background.

The varying string coupling in the NS5-brane solution is

$$e^{2\phi_{loc}(r)} = g_s^2 + \sum_{i=1}^{n_5} \frac{\alpha'}{2\pi^2(r - r_i)^2} \quad (3.41)$$

where r is the radial coordinate, the transverse string frame metric being

$$e^{2\phi_{loc}(r)}(dr^2 + r^2 d\Omega^2). \quad (3.42)$$

To check if the NS5 branes can significantly change the string coupling – and hence KK5 tension – let us start from the previous results and estimate the magnitude of the effect in that configuration. From (3.41) and the result (3.32) that without the present effect, $g_s \sim 1/K^{1/2} \sim 1/M^{1/2}$, we see that within a radial position r_* of order $n_5^{1/2} M^{1/2}$, the string coupling is affected significantly by the NS5-branes. Moreover, the metric (3.42) at most increases the minimal length of the path traced by the fivebrane in the u, \tilde{u} directions by one power of $e^{\phi_{loc}}$, which cannot cancel the effect of the tension (which is quadratic, of order $e^{-2\phi_{loc}}$). Also, the NS5-brane solution does not warp the string frame metric in the directions along its worldvolume, including x, \tilde{x} . Since the KK5-brane core and L_u are of size $M^{1/2}$ in the original solution, we see that introducing any NS5-branes would begin to reduce the tension of the whole KK5-brane set.

Since the KK5-branes are extended in one direction in u, \tilde{u} , it is perhaps more natural to distribute our n_5 NS5-branes along this line, putting them out away from the origin of their moduli space. (This is not crucial to get an effect from them, as we just saw, but it is the most symmetric configuration and one to which the system could settle as it minimizes its energy locally.) Since the relevant scales only overlapped marginally in the

above solution, any reduction of the tension of the KK5-branes is sufficient to push the setup in to a regime where the moduli masses are lighter than all KK masses and where the other related scales discussed above are also separated. Obviously this collection of KK 5-branes and NS 5-branes is rather complicated to analyze in detail, but because they carry different charges we do not expect any catastrophic annihilation mode.

As in the above discussion of KK fivebranes, there are possible topological consistency conditions which may constrain n_5 to be a multiple of M , depending on the application of [24] to the $m_0 B_{x\tilde{x}}$ contribution to the generalized fluxes (2.2). Because of this, we have checked that the case $n_5 = M$ remains consistent with the window (3.27); it contributes a new pair of terms going like L_x^2 and $1/L_x^2$ which can be analyzed in the same way as we did (3.28) above. The case $n_5 = M$ is also consistent with our geometry – even if all NS5-branes were together, the coupling (3.41) does not grow to order 1 until $r \sim n_5^{1/2}$, and hence before taking into account the improvement in the scalings from the varying dilaton, their core would be about the size of the transverse space (as was true for the original KK fivebranes).

3.7. Angular and Fivebrane Moduli

Having addressed the runaway moduli from the volumes and dilaton, let us return to the angular moduli. These fall into two categories:

- (1) Angular moduli sourced by the ingredients listed above
- (2) Angular moduli not sourced by the ingredients listed above.

Those in category (1) must be analyzed on the same footing as the moduli L, L_x, g listed above, to ensure that they do not turn on to large enough values to potentially destabilize the dS minimum. Those in category (2), such as some of RR axions we will discuss, will not destabilize the dS minimum wherever they end up, and will generically be lifted by higher order, lower-scale corrections. Many of the moduli in both categories could be projected out by an orbifold version of the above construction (with appropriate corrections to the order one factors in the volume and potential). However, it is also interesting to pursue their stabilization without using that crutch, since many of the existing ingredients provide the requisite forces in a natural way.

First, recall that the continuous Wilson line moduli $b_{I\tilde{J}}$ coming from the B field (3.19) are lifted by the F_2^2 term, since $F_2 = m_0 B + dC_1$. The metric flux renders the components $B_{u\tilde{x}} = -B_{\tilde{u}x}$ discrete because the x, \tilde{x} circles are torsion 1-cycles. (These are some of the discrete Wilson lines discussed in §3.2.)

Next, consider the metric moduli. As discussed around (3.8), the curvature alone gives positive mass squareds to $\mathcal{G}_{x\tilde{x}}$, $\mathcal{G}_{u_i\tilde{u}_j}$, and $\mathcal{G}_{\tilde{x}\tilde{u}_j} = \mathcal{G}_{xu_j}$. However, we must consider other contributions to their potential from the other ingredients in the construction.

Let us discuss first the angular degree of freedom η defined in (3.23)(3.24). Since this comes from $G_{x\tilde{x}}$, this particular modulus is not easy to project out by an orbifold without also projecting out the $B_{x\tilde{x}}$ contribution which played an important role above. Several of the ingredients we have specified depend on this angle, as explained previously, in the discussion following (3.23)(3.24). Including these effects, we obtain a potential energy of the form

$$\begin{aligned} \frac{\mathcal{U}}{M_4^4} &= ag^2 - bg^3 + cg^4 \\ &= g^2 \left([1 + (\eta - \eta_{q'})^2 + \dots] M^2 \frac{L_x^4}{2L^6} + (4 \frac{n_K}{\sqrt{\eta}}) \frac{L_x^{5/2}}{L^{9/2}} + \frac{3p^2}{2\eta(\alpha')^2 L^6} \right) - g^3 \left(2\sqrt{\frac{2}{\eta\alpha'}} |pm_0| \right) \\ &+ g^4 \alpha' \left(\frac{m_0^2}{4} L^6 + 4\pi^4 m_0^2 \left(\frac{q}{M} \right)^2 \frac{L^6}{L_x^4} + \left(\frac{r}{M} \right)^2 \frac{16\pi^4 m_0^2 L^3}{\eta L_x} + \left(\frac{r}{M} \right)^4 \frac{2^8 \pi^8 m_0^2}{L_x^2} + \frac{K^2}{4L^6(\alpha')^6} \right) \end{aligned} \quad (3.43)$$

Correspondingly, the quantity (3.26) becomes

$$\begin{aligned} \frac{4ac}{b^2} &\equiv 1 + \delta(L, L_x, \eta) \\ &= \frac{\eta(\alpha')^2}{2p^2 m_0^2} \left([1 + (\eta - \eta_{q'})^2 + \dots] M^2 \frac{L_x^4}{2L^6} + (4 \frac{n_K}{\sqrt{\eta}}) \frac{L_x^{5/2}}{L^{9/2}} + \frac{3p^2}{2\eta(\alpha')^2 L^6} \right) \\ &\times \left(\frac{m_0^2}{4} L^6 + 4\pi^4 m_0^2 \left(\frac{q}{M} \right)^2 \frac{L^6}{L_x^4} + \left(\frac{r}{M} \right)^2 \frac{16\pi^4 m_0^2 L^3}{\eta L_x} + \left(\frac{r}{M} \right)^4 \frac{2^8 \pi^8 m_0^2}{L_x^2} + \frac{K^2}{4L^6(\alpha')^6} \right) \end{aligned} \quad (3.44)$$

As it stands, in (3.44) there is a runaway direction in which $\eta \rightarrow 0$ with $L^3 L_x \propto 1/\eta$. (In this limit $\eta \rightarrow 0$, we find also that the curvature potential has a term proportional to $1/\eta$.) But for sufficiently large $\eta_{q'}$, of order 1, the stabilizing mass from the curvature term competes with the tadpoles from the other ingredients; we expect the two to balance to yield a *local* minimum in this direction. There is certainly a local minimum in the η as well as in the g, L, L_x directions, at the point where $\eta \sim 1 \sim \eta_q$ and g, L, L_x sit at the minimum derived in §3.4. The remaining question is whether the other moduli can mix with η significantly enough to introduce large enough off-diagonal contributions to the mass matrix to destabilize the system. Generically, near $\eta = 1$ the second derivatives $L^2 \partial_L^2, L_x^2 \partial_{L_x}^2, g^2 \partial_g^2, \partial_\eta^2$, and the mixed terms $L \partial_L \partial_\eta, \dots$ acting on $(4ac/b^2)$ are of the same order (that of the typical term in $4ac/b^2$). In a similar way to the way we tune the

cosmological constant by picking K/M , we can pick r/M and if necessary the coefficients of additional fluxes we did not use yet (such as other two-form or four-form fluxes) to help tune the off-diagonal elements of the mass matrix to be smaller than the diagonal elements if necessary. The degree to which this extra tuning is required depends on the order one coefficients involved in determining these derivatives, which we have not explicitly calculated here.

Let us now discuss the other angular moduli, which have some similar features. In general, turning on angles in one direction $G_{\mu\nu}$ at fixed volume increases the size of the cycle in either the μ or ν direction. For example, turning on $\mathcal{G}_{u_1\tilde{u}_1}$ at fixed volume of the two-torus in the u_1, \tilde{u}_1 directions increases the size of the cycle wrapped by the O6-plane and the $KK5$ branes, and it decreases the size of the 3-cycle threaded by the H flux. Let $\rho_{1\bar{1}}\sqrt{2}L_{u_1}$ denote the size of the cycle in the $u_1 + \tilde{u}_1$ direction (normalized so that $\rho_{1\bar{1}} = 1$ for the diagonal metric (3.1)). The O6-plane energy and the $KK5$ -brane energy each get multiplied by $\rho_{1\bar{1}}$, and the H flux term gets a factor of $\rho_{1\bar{1}}^2$.

Taking these effects into account for all angles (defining $\rho_{\mu\nu}$ similarly to how we just defined $\rho_{1\bar{1}}$), the quantity $4ac/b^2$ (3.26) takes the form

$$\begin{aligned} \frac{4ac}{b^2} &\equiv 1 + \delta(L, L_x, \rho) \\ &= \frac{(\alpha')^2}{\theta_O^2(\rho_{\mu\nu})(2p^2m_0^2)} \left(\theta_{\mathcal{R}}(\rho)M^2\frac{L_x^4}{2L^6} + \theta_{KK}(\rho_{\mu\nu})(4n_K)\frac{L_x^{5/2}}{L^{9/2}} + \frac{\theta_H(\rho_{\mu\nu})3p^2}{2(\alpha')^2L^6} \right) \\ &\times \left(\frac{m_0^2}{4}L^6 + 4\pi^4m_0^2\left(\frac{q}{M}\right)^2\frac{L^6}{L_x^4} + \theta_{r,2}(\rho)8\pi^4m_0^2\left(\frac{r}{M}\right)^2\frac{L^3}{L_x} + \theta_{r,4}(\rho)\left(\frac{r}{M}\right)^4\frac{2^6\pi^8m_0^2}{L_x^2} + \frac{K^2}{4L^6(\alpha')^6} \right) \end{aligned} \quad (3.45)$$

where we have schematically indicated by functions $\theta(\rho)$ the dependence on $\rho_{\mu\nu}$ of those ingredients which are sensitive to these angles.

There are two types of angles in our problem: those which are sourced by the O6-plane and H flux, analogously to the $G_{x\tilde{x}}$ mode discussed above ($\mathcal{G}_{u_1\tilde{u}_1}$ and $\mathcal{G}_{u_2\tilde{u}_2}$), and those which are not ($\mathcal{G}_{u_1u_2} = \mathcal{G}_{\tilde{u}_1\tilde{u}_2}$, $\mathcal{G}_{u_1\tilde{u}_2} = \mathcal{G}_{\tilde{u}_1u_2}$, and $\mathcal{G}_{x,\tilde{u}_j} = \mathcal{G}_{\tilde{x},u_j}$).

Let us start with the former case. As in our discussion of $G_{x\tilde{x}}$, the mass squared introduced by the curvature has the right shape and order of magnitude to provide a local minimum when balanced against the tadpoles from the O6 and H flux, depending on order one coefficients and on discrete parameters that can be tuned. We have not analyzed these coefficients in detail, so let us mention two other methods for stabilizing $\rho_{u_1\tilde{u}_1}, \rho_{u_2,\tilde{u}_2}$. First, we can consider adding an additional sector of KK fivebranes to avoid

the $\rho \rightarrow \infty$ direction. Namely, a KK5-brane wrapped on the x, \tilde{x} direction, and with fiber circle $u_1 + \tilde{u}_1 + u_2 + \tilde{u}_2$ gives a potential term scaling like $g^2 \rho^2 L_x / L^3$. This prevents the potential runaway limit of (3.45) to large ρ , and the new ingredient is at most of the same order as the previous contributions to the potential, and so can be added consistently. Alternatively, it is possible to enforce $\rho = 1$ by an orbifold. One example is a \mathbb{Z}_2 orbifold under which $u_j \rightarrow -u_j, \tilde{u}_j \rightarrow -\tilde{u}_j, x \rightarrow x, \tilde{x} \rightarrow \tilde{x}$ (either combined with a transverse shift to remove twisted moduli, or with extra fivebranes wrapped around the blowup cycles, carrying worldvolume gauge flux to stabilize their sizes). This introduces an O6 fixed plane in the directions $u_j - \tilde{u}_j, x + \tilde{x}$, whose linear potential for $\mathcal{G}_{u_1 \tilde{u}_1}$ and $\mathcal{G}_{u_2 \tilde{u}_2}$ cancels against that of the original O6-plane in the symmetric configuration (3.1), making it manifest that the curvature mass term [28] suffices to lift this angular direction. (Various other factors in the tadpole cancellation condition and in the potential also change in the orbifold case, and the contributions proportional to r/M – which were not crucial in the stabilization above – are projected out. A preliminary check of the coefficients resulted in parameters still consistent with the window (3.27); it would be useful to systematically study different orbifold space groups with regard to the question of the constant terms in (3.26) resulting from the normalized volume and tadpole cancelation conditions.)

The angular moduli which are not sourced by the O6-plane and H flux are also stabilized by curvature in the case $\mathcal{G}_{u_1 \tilde{u}_2} = \mathcal{G}_{\tilde{u}_1 u_2}$. In the case $\mathcal{G}_{u_1 u_2} = \mathcal{G}_{\tilde{u}_1 \tilde{u}_2}$ they are stabilized by the KK fivebranes: turning on angles between the u_1, \tilde{u}_1 and u_2, \tilde{u}_2 directions at fixed volume increases the lengths of the cycles the KK5-branes are wrapped on, while not affecting the volume of the cycle wrapped by the O6-plane or the cycle threaded by the H flux. (A similar mechanism arises in intersecting brane models [51].)

This leaves us with $\mathcal{G}_{u_j \tilde{x}} = \mathcal{G}_{x \tilde{u}_j}$ (if the \mathbb{Z}_2 orbifold option described above is not taken – this would simply project them out). These are metric Wilson lines. Recall that the metric flux lifts the $B_{u\tilde{x}} = -B_{\tilde{u}x}$ modes. H flux similarly lifts the modes $\mathcal{G}_{u_j \tilde{x}} = \mathcal{G}_{x \tilde{u}_j}$, by T -duality. In our case, the H flux potential is minimized at one of two values, since we took the H flux quantum number to be 2. That is, these metric Wilson lines are discrete. These modes are thus similar to those discussed above, but with the positive mass squared contribution coming from H flux rather than from the curvature term.

The RR axions are either fixed by virtue of the Chern-Simons terms in (2.2), or contribute subdominantly to the potential (hence falling in category (2) above). There are

three components of C_3 invariant under the orientifold action (3.10). They are dual to the components of H indicated in (3.13), and are stabilized by the H flux, since

$$\tilde{F}_6 = dC_5 - C_3 \wedge H_3 + \frac{m_0}{6} B \wedge B \wedge B \quad (3.46)$$

Similar comments apply to C_1 and C_5 : components which are not projected out by the orientifold action or lifted by the $|\tilde{F}_4|^2$ terms are fixed by higher order effects which generate the axion potential.

Finally, let us return to the motion moduli of the KK fivebranes. As discussed above, the discrete torsion requires the M branes to sit together, projecting out their relative motions. Given the mechanism suggested in §3.6, the size of the set of $n_K = M$ KK fivebranes is parametrically smaller than L_u , and in all versions of the construction it is no larger than the transverse space. The overall position of the KK5-branes is inconsequential to the stabilization of the runaway moduli, and hence directions which are not fixed by the curvature potential are in category (2) above; again, if we invoke the method of §3.6, then the position of the KK 5-branes is localized near the source of enhanced $e^{\phi_{loc}}$. There are no isometries in the transverse directions to the KK fivebranes, so in any case their position moduli will be lifted by effects to do with the ambient curvature of the space transverse to the KK5-branes.

3.8. Numerical solution

We showed analytically above that a local minimum of (3.25) exists for appropriate choices of discrete quantum numbers. We have checked this numerically using Mathematica. In doing so it was again useful to follow the procedure used above, first finding a minimum of $4ac/b^2$ at some $L = L_0, L_x = L_{x0}$ (tuning the discrete quantum numbers f_6, M to arrange for the minimum value δ_0 of $1 - 4ac/b^2$ to be close to 0). Next, we minimized $\mathcal{U}(g, L_0, L_{x0})$ with respect to g . Then searching near that point for a minimum in all directions yields the expected solution. As a specific example, with $n_K = M = 10$, $f_6 = 80$, $q = 1$, and $r = 1$ the minimum of $4ac/b^2$ is at approximately 1.0003 (so $\delta_0 = 0003 \ll 1$ is very small, putting us in the regime of applicability of the analysis in §2). The potential is minimized at $\mathcal{U}/M_4^2 \approx 10^{-13}$ with $g \approx 0.00015$, $L \approx 15.3$, $L_x \approx 2.1$.

Note that in this solution, with M taken to be 10, L_x is *not* substringy in size, the string coupling g_s is of order 1/2, and the T-dual string coupling is also not strong. Scaling M up pushes L_x down as discussed in the text, but at modest finite values of the parameters such as those given here one can obtain less extreme behavior.

Plotting the potential \mathcal{U}/M_4^4 in the g direction yields

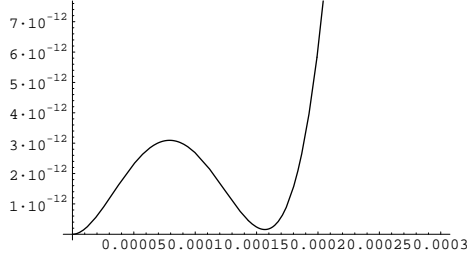


fig. 1

and in the L and L_x directions we obtain (note that the horizontal axis is not placed at zero)

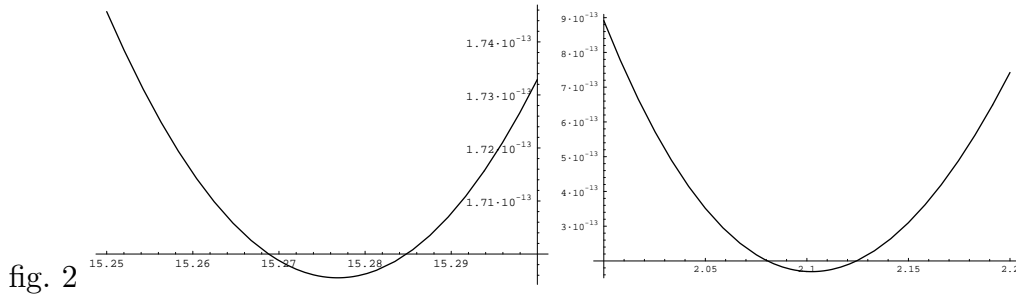


fig. 2

Note that this numerical analysis does not explicitly include the angular variables discussed in §3.7.

3.9. Metastability

As emphasized in [11,4], models of de Sitter which are obtained in a weak coupling regime are only metastable.¹⁰ In the present case, the decay yields a ten-dimensional generalization of a Bianchi cosmology, with different directions evolving anisotropically – the x, \tilde{x} directions shrinking and the others expanding. As discussed above, a nilmanifold in vacuum can be T-dualized to an expanding space with H flux [32], but since the $B_{x\tilde{x}}$ field is nontrivial in our solution the element of the T-duality group which is relevant at a given radius is not the simple one considered in [32]. The question of whether one can or cannot T-dualize to large radius, perhaps via a time-dependent T-duality cascade¹¹, is an interesting one.

¹⁰ For a recent comparative study of decays in a subset of string theoretic dS models, see for example [52].

¹¹ I thank S. Kachru for this suggestion. These questions are also related to those analyzed in [53].

4. Toward Simple de Sitter Solutions

We expect similar solutions from a compactification on a product of two Sol 3-manifolds; let us sketch the analogues of the steps given above for the nilmanifold case. The sol 3-geometry is

$$\begin{aligned} ds^2 &= \alpha' (L_y^2(e^{2z}dy_1^2 + e^{-2z}dy_2^2) + L_z^2dz^2) \\ &= \alpha' (L_y^2\omega_1^2 + L_y^2\omega_2^2 + L_z^2\omega_3^2) \end{aligned} \tag{4.1}$$

where $\omega_1 = L_y e^z dy_1$, $\omega_2 = L_y e^{-z} dy_2$, and $\omega_3 = L_z dz$. This geometry has three independent isometries consisting of shifts of y_1, y_2 and shifts of z combined with rescalings of y_1, y_2 .

Compact solmanifolds \mathcal{S}_3 are obtained by projecting (4.1) by a discrete subgroup of the isometry group. As one moves around the z direction, the τ parameter of the T^2 in the y_1, y_2 directions undergoes an $SL(2, \mathbb{Z})$ transformation. This is analogous to (3.2) except that here the $SL(2, \mathbb{Z})$ transformation must be more general than $\tau \rightarrow \tau + 1$ in order to provide a consistent compactification; it shrinks the torus exponentially in the y_1 direction and expands it exponentially in the y_2 direction. (Related to this, the solmanifold has a rich fundamental group, which is of exponential growth.)

As with the nilmanifold, the sol manifold has fewer massless moduli than the corresponding T^3 : the relations

$$d\omega_1 = \omega_3 \wedge \omega_1 \quad d\omega_2 = -\omega_3 \wedge \omega_2 \tag{4.2}$$

mean that the homology groups H^1 and H^2 are each reduced by two dimensions as compared to a torus.

The scalar curvature of \mathcal{S}_3 is $-2/L_z^2 = -2L_y^4/L^6$ where $L^6 = L_y^4 L_z^2$. Upon compactification to four dimensions on a product of two solmanifolds this will lead to a positive curvature potential analogous to (3.7), with discrete parameters analogous to M in (3.7) which have to do with the choice of $SL(2, \mathbb{Z})$ element used in the compactification procedure. As in the case of the nilmanifold, the rich topology of \mathcal{S}_3 provides a place for branes to wrap and a potential source of fractional Chern-Simons invariants. All this again leads to a four dimensional scalar potential analogous to (3.25), with the role of x played by y_1, y_2 and the role of u_1, u_1 played by z . It would be interesting to flesh this out explicitly to see if again the discrete parameters available are sufficient to tune the system into the range (2.9).

5. Discussion

In this work, we proposed a relatively simple and explicit class of de Sitter models in string theory. We showed how a few ingredients suffice to produce a potential for moduli which exhibits metastable minima at positive vacuum energy, seven independent terms being involved in the basic stabilization of g, L , and L_x . Clearly an important direction for further work is fleshing out further the methods in §3.6 and §3.7 for stabilizing the angular moduli and for separating the scales. A convenient feature of the background is its weak curvature and 10d string coupling, and its 10d supersymmetry, which make possible a controlled analysis of the moduli and the KK and winding modes. On the other hand, the SUSY breaking effects of the curvature and KK fivebrane configuration facilitate moduli stabilization by introducing useful competing forces (which would vanish in the lower-energy SUSY models based on Calabi-Yau manifolds with the subset of ingredients analyzed in the no-go theorem of [33]).

One of the main general lessons is that “metric flux” and wrapped KK 5-branes yield forces whose dependence on the moduli is appropriate to “uplift” the potential for the runaway moduli g, L, L_x in the type IIA AdS_4 solutions of the sort studied in [14]. The most complicated aspect of the specific models is probably the 5-brane dynamics.

One natural question is whether a version of this mechanism exists in which lower energy supersymmetry is preserved. In [24] it was suggested that various “metric fluxes” could cancel the charges of KK fivebranes as well as $NS5$ -branes. Such a construction could be analogous to the way fluxes cancel the $O3$ -plane charges in the type IIB models of [10,4]. Combining this idea of [24] with the mechanism described here might be a concrete place to seek models with lower scale supersymmetry breaking.

It would be interesting to apply our construction to the problem of explicitly modeling inflation in string theory (for recent reviews see [2]). One question is whether our fractional Chern-Simons invariants could also help tune the inflationary ϵ and η parameters to be small. It might also be interesting to introduce particle physics sectors to these models, perhaps using stretched D-branes within the bulwark of KK5-branes to form brane constructions of the relevant field theories.

Explicit models of de Sitter (and also anti de Sitter, if we reduce K/M) may facilitate the derivation of concrete holographic duals. Some progress toward a general framework for duals of metastable de Sitter space have appear in [54]. Ideas for unveiling the degrees of freedom of the dual out on its approximate moduli space can be found in [55], by trading

the flux for branes in analogy with the construction of the Coulomb branch in more familiar versions of the AdS/CFT correspondence. In the present case, one might trade the H field and metric flux for $NS5$ -branes and KK monopoles.

Acknowledgements

It is a pleasure to thank B. Florea, T. Grimm, and A. Tomasiello for very useful discussions; O. Aharony, D. Green, S. Kachru, L. McAllister, W. Taylor and A. Westphal for useful discussions as well as very useful comments on a draft; and A. Lawrence, A. Maloney and D. Starr for ongoing discussions on the “fundamental” physics of compact target spaces and their T duals. This research is supported in part by NSF grant PHY-0244728, and in part by the DOE under contract DE-AC03-76SF00515.

References

- [1] A. R. Frey, “Warped strings: Self-dual flux and contemporary compactifications,” arXiv:hep-th/0308156; E. Silverstein, “TASI / PiTP / ISS lectures on moduli and microphysics,” arXiv:hep-th/0405068; M. Grana, “Flux compactifications in string theory: A comprehensive review,” Phys. Rept. **423**, 91 (2006) [arXiv:hep-th/0509003]; J. Polchinski, “The cosmological constant and the string landscape,” arXiv:hep-th/0603249; M. R. Douglas and S. Kachru, “Flux compactification,” Rev. Mod. Phys. **79**, 733 (2007) [arXiv:hep-th/0610102]; R. Bousso, “TASI Lectures on the Cosmological Constant,” arXiv:0708.4231 [hep-th].
- [2] S. H. Henry Tye, “Brane inflation: String theory viewed from the cosmos,” arXiv:hep-th/0610221; J. M. Cline, “String cosmology,” arXiv:hep-th/0612129; R. Kallosh, “On Inflation in String Theory,” arXiv:hep-th/0702059; C. P. Burgess, “Lectures on Cosmic Inflation and its Potential Stringy Realizations,” arXiv:0708.2865 [hep-th]; L. McAllister and E. Silverstein, “String Cosmology: A Review,” arXiv:0710.2951 [hep-th].
- [3] A. Maloney, E. Silverstein and A. Strominger, “De Sitter space in noncritical string theory,” Cambridge 2002, The future of theoretical physics and cosmology* 570-591 [arXiv:hep-th/0205316].
- [4] S. Kachru, R. Kallosh, A. Linde and S. P. Trivedi, “De Sitter vacua in string theory,” Phys. Rev. D **68**, 046005 (2003) [arXiv:hep-th/0301240].
- [5] V. Balasubramanian, P. Berglund, J. P. Conlon and F. Quevedo, “Systematics of moduli stabilisation in Calabi-Yau flux compactifications,” JHEP **0503**, 007 (2005) [arXiv:hep-th/0502058]; A. Saltman and E. Silverstein, “The scaling of the no-scale potential and de Sitter model building,” JHEP **0411**, 066 (2004) [arXiv:hep-th/0402135]; A. Misra and P. Shukla, “Area Codes, Large Volume (Non-)Perturbative α' - and Instanton - Corrected Non-supersymmetric (A)dS minimum, the Inverse Problem and Fake Superpotentials for Multiple-Singular-Loci-Two-Parameter Calabi-Yau’s,” arXiv:0707.0105 [hep-th].
- [6] A. Saltman and E. Silverstein, “A new handle on de Sitter compactifications,” JHEP **0601**, 139 (2006) [arXiv:hep-th/0411271].
- [7] K. Becker and M. Becker, “M-Theory on Eight-Manifolds,” Nucl. Phys. B **477**, 155 (1996) [arXiv:hep-th/9605053]; K. Becker and M. Becker, “Supersymmetry breaking, M-theory and fluxes,” JHEP **0107**, 038 (2001) [arXiv:hep-th/0107044].
- [8] S. Gukov, C. Vafa and E. Witten, “CFT’s from Calabi-Yau four-folds,” Nucl. Phys. B **584**, 69 (2000) [Erratum-ibid. B **608**, 477 (2001)] [arXiv:hep-th/9906070].
- [9] K. Dasgupta, G. Rajesh and S. Sethi, “M theory, orientifolds and G-flux,” JHEP **9908**, 023 (1999) [arXiv:hep-th/9908088].
- [10] S. B. Giddings, S. Kachru and J. Polchinski, “Hierarchies from fluxes in string compactifications,” Phys. Rev. D **66**, 106006 (2002) [arXiv:hep-th/0105097].

- [11] E. Silverstein, “(A)dS backgrounds from asymmetric orientifolds,” arXiv:hep-th/0106209.
- [12] R. Bousso and J. Polchinski, “Quantization of four-form fluxes and dynamical neutralization of the cosmological constant,” JHEP **0006**, 006 (2000) [arXiv:hep-th/0004134].
- [13] J. L. Feng, J. March-Russell, S. Sethi and F. Wilczek, “Saltatory relaxation of the cosmological constant,” Nucl. Phys. B **602**, 307 (2001) [arXiv:hep-th/0005276].
- [14] O. DeWolfe, A. Giryavets, S. Kachru and W. Taylor, “Type IIA moduli stabilization,” JHEP **0507**, 066 (2005) [arXiv:hep-th/0505160].
- [15] T. W. Grimm and J. Louis, “The effective action of type IIA Calabi-Yau orientifolds,” Nucl. Phys. B **718**, 153 (2005) [arXiv:hep-th/0412277].
- [16] G. Villadoro and F. Zwirner, “ $N = 1$ effective potential from dual type-IIA D6/O6 orientifolds with general fluxes,” JHEP **0506**, 047 (2005) [arXiv:hep-th/0503169].
- [17] T. Banks and K. van den Broek, “Massive IIA flux compactifications and U-dualities,” JHEP **0703**, 068 (2007) [arXiv:hep-th/0611185].
- [18] F. Denef, M. R. Douglas, B. Florea, A. Grassi and S. Kachru, “Fixing all moduli in a simple F-theory compactification,” Adv. Theor. Math. Phys. **9**, 861 (2005) [arXiv:hep-th/0503124].
- [19] F. Denef, M. R. Douglas and B. Florea, “Building a better racetrack,” JHEP **0406**, 034 (2004) [arXiv:hep-th/0404257].
- [20] J. P. Conlon and F. Quevedo, “On the explicit construction and statistics of Calabi-Yau flux vacua,” JHEP **0410**, 039 (2004) [arXiv:hep-th/0409215].
- [21] T. House and E. Palti, “Effective action of (massive) IIA on manifolds with $SU(3)$ structure,” Phys. Rev. D **72**, 026004 (2005) [arXiv:hep-th/0505177].
- [22] P. G. Camara, A. Font and L. E. Ibanez, “Fluxes, moduli fixing and MSSM-like vacua in a simple IIA orientifold,” JHEP **0509**, 013 (2005) [arXiv:hep-th/0506066].
- [23] D. Lust and D. Tsimpis, “Supersymmetric AdS(4) compactifications of IIA supergravity,” JHEP **0502**, 027 (2005) [arXiv:hep-th/0412250]; M. Grana, R. Minasian, M. Petrini and A. Tomasiello, “A scan for new $N=1$ vacua on twisted tori,” JHEP **0705**, 031 (2007) [arXiv:hep-th/0609124].
- [24] G. Villadoro and F. Zwirner, “On general flux backgrounds with localized sources,” arXiv:0710.2551 [hep-th]; G. Villadoro and F. Zwirner, “Beyond Twisted Tori,” Phys. Lett. B **652**, 118 (2007) [arXiv:0706.3049 [hep-th]].
- [25] J. Shelton, W. Taylor and B. Wecht, “Generalized flux vacua,” JHEP **0702**, 095 (2007) [arXiv:hep-th/0607015].
- [26] D. Robbins and T. Wrase, “D-Terms from Generalized NS-NS Fluxes in Type II,” arXiv:0709.2186 [hep-th]. M. Ihl, D. Robbins and T. Wrase, “Toroidal Orientifolds in IIA with General NS-NS Fluxes,” JHEP **0708**, 043 (2007) [arXiv:0705.3410 [hep-th]]. M. Ihl and T. Wrase, “Towards a realistic type IIA $T^{*6}/Z(4)$ orientifold model with background fluxes. I: Moduli stabilization,” JHEP **0607**, 027 (2006) [arXiv:hep-th/0604087].

- [27] A. Lawrence, T. Sander, M. B. Schulz and B. Wecht, “Torsion and Supersymmetry Breaking,” arXiv:0711.4787 [hep-th].
- [28] J. Scherk and J. H. Schwarz, “How To Get Masses From Extra Dimensions,” Nucl. Phys. B **153**, 61 (1979); N. Kaloper and R. C. Myers, “The O(dd) story of massive supergravity,” JHEP **9905**, 010 (1999) [arXiv:hep-th/9901045]. C. M. Hull and R. A. Reid-Edwards, “Flux compactifications of string theory on twisted tori,” arXiv:hep-th/0503114.
- [29] S. Gukov, S. Kachru, X. Liu and L. McAllister, “Heterotic moduli stabilization with fractional Chern-Simons invariants,” Phys. Rev. D **69**, 086008 (2004) [arXiv:hep-th/0310159].
- [30] M. Dine, A. Morisse, A. Shomer and Z. Sun, “IIA moduli stabilization with badly broken supersymmetry,” arXiv:hep-th/0612189.
- [31] D. Robbins and T. Wrase, “D-Terms from Generalized NS-NS Fluxes in Type II,” arXiv:0709.2186 [hep-th].
- [32] S. Kachru, M. B. Schulz, P. K. Tripathy and S. P. Trivedi, “New supersymmetric string compactifications,” JHEP **0303**, 061 (2003) [arXiv:hep-th/0211182].
- [33] M. P. Hertzberg, S. Kachru, W. Taylor and M. Tegmark, “Inflationary Constraints on Type IIA String Theory,” arXiv:0711.2512 [hep-th]; M. P. Hertzberg, M. Tegmark, S. Kachru, J. Shelton and O. Ozcan, “Searching for Inflation in Simple String Theory Models: An Astrophysical Perspective,” arXiv:0709.0002 [astro-ph].
- [34] F. Saueressig, U. Theis and S. Vandoren, “On de Sitter vacua in type IIA orientifold compactifications,” Phys. Lett. B **633**, 125 (2006) [arXiv:hep-th/0506181].
- [35] M. Gromov, “Groups of Polynomial growth and Expanding Maps”, Publications mathématiques I.H..S., 53, 1981
- [36] D. Green, A. Lawrence, J. McGreevy, D. R. Morrison and E. Silverstein, “Dimensional Duality,” Phys. Rev. D **76**, 066004 (2007) [arXiv:0705.0550 [hep-th]]. J. McGreevy, E. Silverstein and D. Starr, “New dimensions for wound strings: The modular transformation of geometry to topology,” Phys. Rev. D **75**, 044025 (2007) [arXiv:hep-th/0612121]. E. Silverstein, “Dimensional mutation and spacelike singularities,” Phys. Rev. D **73**, 086004 (2006) [arXiv:hep-th/0510044].
- [37] A. Maloney, unpublished; A. Lawrence, A. Maloney, E. Silverstein, and D. Starr, works in progress.
- [38] G. W. Gibbons and S. W. Hawking, “Cosmological Event Horizons, Thermodynamics, And Particle Creation,” Phys. Rev. D **15**, 2738 (1977).
- [39] J. Polchinski, “String theory. Vol. 2: Superstring theory and beyond,” *Cambridge, UK: Univ. Pr. (1998) 531 p*
- [40] A. Adams, M. Ernebjerg and J. M. Lapan, “Linear models for flux vacua,” arXiv:hep-th/0611084.

- [41] A. R. Frey and J. Polchinski, “ $N = 3$ warped compactifications,” *Phys. Rev. D* **65**, 126009 (2002) [arXiv:hep-th/0201029].
- [42] M. B. Green, J. H. Schwarz and E. Witten, *Cambridge, U.K.: Univ. Pr. (1987) 596 P. (Cambridge Monographs On Mathematical Physics)*
- [43] T. Ishii, G. Ishiki, S. Shimasaki and A. Tsuchiya, “T-duality, fiber bundles and matrices,” *JHEP* **0705**, 014 (2007) [arXiv:hep-th/0703021].
- [44] C. Vafa, “Modular Invariance And Discrete Torsion On Orbifolds,” *Nucl. Phys. B* **273**, 592 (1986).
- [45] E. R. Sharpe, “Recent developments in discrete torsion,” *Phys. Lett. B* **498**, 104 (2001) [arXiv:hep-th/0008191].
- [46] D. J. Gross and M. J. Perry, “Magnetic Monopoles In Kaluza-Klein Theories,” *Nucl. Phys. B* **226**, 29 (1983); R. d. Sorkin, “Kaluza-Klein Monopole,” *Phys. Rev. Lett.* **51**, 87 (1983).
- [47] R. Gregory, J. A. Harvey and G. W. Moore, “Unwinding strings and T-duality of Kaluza-Klein and H-monopoles,” *Adv. Theor. Math. Phys.* **1**, 283 (1997) [arXiv:hep-th/9708086]; D. Tong, “NS5-branes, T-duality and worldsheet instantons,” *JHEP* **0207**, 013 (2002) [arXiv:hep-th/0204186]; H. Ooguri and C. Vafa, “Two-Dimensional Black Hole and Singularities of CY Manifolds,” *Nucl. Phys. B* **463**, 55 (1996) [arXiv:hep-th/9511164].
- [48] See e.g. G. W. Moore and E. Witten, “Self-duality, Ramond-Ramond fields, and K-theory,” *JHEP* **0005**, 032 (2000) [arXiv:hep-th/9912279]; A. M. Uranga, “D-brane probes, RR tadpole cancellation and K-theory charge,” *Nucl. Phys. B* **598**, 225 (2001) [arXiv:hep-th/0011048].
- [49] R. Minasian and G. W. Moore, “K-theory and Ramond-Ramond charge,” *JHEP* **9711**, 002 (1997) [arXiv:hep-th/9710230]; E. Witten, “D-branes and K-theory,” *JHEP* **9812**, 019 (1998) [arXiv:hep-th/9810188]. D. S. Freed and E. Witten, “Anomalies in string theory with D-branes,” arXiv:hep-th/9907189; R. Minasian and G. W. Moore, “K-theory and Ramond-Ramond charge,” *JHEP* **9711**, 002 (1997) [arXiv:hep-th/9710230].
- [50] C. Vafa and E. Witten, “On orbifolds with discrete torsion,” *J. Geom. Phys.* **15**, 189 (1995) [arXiv:hep-th/9409188].
- [51] R. Blumenhagen, M. Cvetič, P. Langacker and G. Shiu, “Toward realistic intersecting D-brane models,” *Ann. Rev. Nucl. Part. Sci.* **55**, 71 (2005) [arXiv:hep-th/0502005].
- [52] A. Westphal, “Lifetime of Stringy de Sitter Vacua,” arXiv:0705.1557 [hep-th].
- [53] T. Banks, W. Fischler and L. Motl, “Dualities versus singularities,” *JHEP* **9901**, 019 (1999) [arXiv:hep-th/9811194].
- [54] A. Strominger, “Inflation and the dS/CFT correspondence,” *JHEP* **0111**, 049 (2001) [arXiv:hep-th/0110087]. T. Banks and W. Fischler, “An holographic cosmology,” arXiv:hep-th/0111142. M. Alishahiha, A. Karch, E. Silverstein and D. Tong, “The

- dS/dS correspondence,” AIP Conf. Proc. **743**, 393 (2005) [arXiv:hep-th/0407125].
M. Alishahiha, A. Karch and E. Silverstein, “Hologravity,” JHEP **0506**, 028 (2005) [arXiv:hep-th/0504056].
B. Freivogel, V. E. Hubeny, A. Maloney, R. Myers, M. Rangamani and S. Shenker, “Inflation in AdS/CFT,” JHEP **0603**, 007 (2006) [arXiv:hep-th/0510046].
B. Freivogel, Y. Sekino, L. Susskind and C. P. Yeh, “A holographic framework for eternal inflation,” Phys. Rev. D **74**, 086003 (2006) [arXiv:hep-th/0606204].
N. Arkani-Hamed, S. Dubovsky, A. Nicolis, E. Trincherini and G. Villadoro, “A Measure of de Sitter Entropy and Eternal Inflation,” JHEP **0705**, 055 (2007) [arXiv:0704.1814 [hep-th]].
- [55] O. Aharony, Y. Antebi and M. Berkooz, to appear; E. Silverstein, “AdS and dS entropy from string junctions or the function of junction conjunctions,” arXiv:hep-th/0308175.

Quick Start Guide

1. Form a Nilmanifold

Twist the two 2-tori around the two circles, deforming them by τ to $\tau+M$ as you go around the circles. Your nilmanifold will try to relax toward zero curvature. Hold it in place while you install the other components.



2. Install KK monopoles stretched along the u directions as described in the manual.

3. Install O-plane and H flux

Carefully fold your manifold, identifying your two Nil 3-manifolds, while **SIMULTANEOUSLY** introducing h_3 units of H flux and f_0 of zero



Warning: Ignoring the H flux would violate Gauss' Law.

4. Install discrete Wilson lines

Thread the product of the two Z_M torsion circles with q/M units of B

5. Introduce Fluxes

Thread your manifold with 6-form flux

6. Relax to minimum

Holding the angular directions fixed, gently release the g , Lu_1, Lu_2 , L_x moduli. Your model should relax to a local minimum in these directions. Next, release the angular modes; check if they relax to a minimum. If not, then add further 5-branes and/or orbifold as described in the manual (section 3.7).